

MICROCOPY RESOLUTION TEST CHART NATIONAL BUREAU OF STANJAMEN SHAPE A

17.5 144

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PLASMA FORMULARY

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NUMERICAL AND ALGEBRAIC

Gain in decibels of P_2 relative to P_1

$$G = 10 \log_{10}(P_2/P_1).$$

To within two percent

$$(2\pi)^{1/2} \approx 2.5$$
; $\pi^2 \approx 10$; $e^3 \approx 20$; $2^{10} \approx 10^3$.

Euler-Mascheroni constant¹ $\gamma = 0.57722$

Gamma Function $\Gamma(x+1) = x\Gamma(x)$:

$$\begin{array}{lll} \Gamma(1/6) = 5.5663 & \Gamma(3/5) = 1.4892 \\ \Gamma(1/5) = 4.5908 & \Gamma(2/3) = 1.3541 \\ \Gamma(1/4) = 3.6256 & \Gamma(3/4) = 1.2254 \\ \Gamma(1/3) = 2.6789 & \Gamma(4/5) = 1.1642 \\ \Gamma(2/5) = 2.2182 & \Gamma(5/6) = 1.1288 \\ \Gamma(1/2) = 1.7725 = \sqrt{\pi} & \Gamma(1) = 1.0 \end{array}$$

Binomial Theorem (good for |x| < 1 or $\alpha = positive integer)$:

$$(1+x)^{\alpha} = \sum_{k=0}^{\infty} {\alpha \choose k} x^{k} \equiv 1 + \alpha x + \frac{\alpha(\alpha-1)}{2!} x^{2} + \frac{\alpha(\alpha-1)(\alpha-2)}{3!} x^{3} + \cdots$$

Rothe-Hagen identity² (good for all complex x, y, z except when singular):

$$\sum_{k=0}^{n} \frac{x}{x+kz} {x+kz \choose k} \frac{y}{y+(n-k)z} {y+(n-k)z \choose n-k}$$

$$= \frac{x+y}{x+y+nz} {x+y+nz \choose n}.$$

Newberger's summation formula³ [good for μ nonintegral, Re $(\alpha+\beta) > -1$]:

$$\sum_{n=-\infty}^{\infty} \frac{(-1)^n J_{\alpha-\gamma n}(z) J_{\beta+\gamma n}(z)}{n+\mu} = \frac{\pi}{\sin \mu \pi} J_{\alpha+\gamma \mu}(z) J_{\beta-\gamma \mu}(z).$$

VECTOR IDENTITIES⁴

Notation: f, g, are scalars: A. B. etc., are vectors: T is a tensor: I is the unit dyad.

(1)
$$A \cdot B \times C = A \times B \cdot C = B \cdot C \times A = B \times C \cdot A = C \cdot A \times B = C \times A \cdot B$$

(2)
$$A \times (B \times C) = (C \times B) \times A = (A \cdot C)B - (A \cdot B)C$$

(3)
$$A \times (B \times C) + B \times (C \times A) + C \times (A \times B) = 0$$

(4)
$$(A \times B) \cdot (C \times D) = (A \cdot C)(B \cdot D) - (A \cdot D)(B \cdot C)$$

(5)
$$(A \times B) \times (C \times D) = (A \times B \cdot D)C - (A \times B \cdot C)D$$

(6)
$$\nabla (fg) = \nabla (gf) = f \nabla g + g \nabla f$$

(7)
$$\nabla \cdot (fA) = f \nabla \cdot A + A \cdot \nabla f$$

(8)
$$\nabla \times (f\mathbf{A}) = f\nabla \times \mathbf{A} + \nabla f \times \mathbf{A}$$

(9)
$$\nabla \cdot (\mathbf{A} \times \mathbf{B}) = \mathbf{B} \cdot \nabla \times \mathbf{A} - \mathbf{A} \cdot \nabla \times \mathbf{B}$$

(10)
$$\nabla \times (\mathbf{A} \times \mathbf{B}) = \mathbf{A}(\nabla \cdot \mathbf{B}) - \mathbf{B}(\nabla \cdot \mathbf{A}) + (\mathbf{B} \cdot \nabla)\mathbf{A} - (\mathbf{A} \cdot \nabla)\mathbf{B}$$

(11)
$$\mathbf{A} \times (\nabla \times \mathbf{B}) = (\nabla \mathbf{B}) \cdot \mathbf{A} - (\mathbf{A} \cdot \nabla) \mathbf{B}$$

(12)
$$\nabla (\mathbf{A} \cdot \mathbf{B}) = \mathbf{A} \times (\nabla \times \mathbf{B}) + \mathbf{B} \times (\nabla \times \mathbf{A}) + (\mathbf{A} \cdot \nabla)\mathbf{B} + (\mathbf{B} \cdot \nabla)\mathbf{A}$$

$$(13) \nabla^2 f = \nabla \cdot \nabla f$$

(14)
$$\nabla^2 \mathbf{A} = \nabla(\nabla \cdot \mathbf{A}) - \nabla \times \nabla \times \mathbf{A}$$

(15)
$$\nabla \times \nabla f = 0$$

(16)
$$\nabla \cdot \nabla \times \mathbf{A} = 0$$

If e_1 , e_2 , e_3 are orthonormal unit vectors, a second-order tensor T can be written in the dyadic form

$$(17) T = \sum_{i,j} T_{ij} e_i e_j$$

In cartesian coordinates the divergence of a tensor is a vector with components

(18)
$$(\nabla \cdot T)_i = \sum_j (\partial T_{ji}/\partial x_j)$$

[This definition is required for consistency with Eq. (29)]. In general

(19)
$$\nabla \cdot (\mathbf{A}\mathbf{B}) = (\nabla \cdot \mathbf{A})\mathbf{B} + (\mathbf{A} \cdot \nabla)\mathbf{B}$$

(20)
$$\nabla \cdot (fT) = \nabla f \cdot T + f \nabla \cdot T$$

Let r = iz + jy + kz be the radius vector of magnitude r, from the origin to the point x, y, z. Then

(21)
$$\nabla \cdot \mathbf{r} = 3$$

(22)
$$\nabla \times \mathbf{r} = \mathbf{0}$$

(23)
$$\nabla r = \mathbf{r}/r$$

$$(24) \nabla (1/r) = -r/r^3$$

$$(25) \nabla \cdot (\mathbf{r}/r^3) = 4\pi \delta(\mathbf{r})$$

(26)
$$\nabla \mathbf{r} = I$$

If V is a volume enclosed by a surface S and dS = ndS, where n is the unit normal outward from V,

$$(27) \int_{V} dV \nabla f = \int_{S} dS f$$

$$(28) \int_{V} dV \nabla \cdot \mathbf{A} = \int_{S} d\mathbf{S} \cdot \mathbf{A}$$

$$(29) \int_{V} dV \nabla \cdot T = \int_{S} d\mathbf{S} \cdot T$$

(30)
$$\int_{V} dV \nabla \times \mathbf{A} = \int_{S} d\mathbf{S} \times \mathbf{A}$$

(31)
$$\int_{V} dV (f \nabla^{2} g - g \nabla^{2} f) = \int_{S} dS \cdot (f \nabla g - g \nabla f)$$

(32)
$$\int_{V} dV (\mathbf{A} \cdot \nabla \times \nabla \times \mathbf{B} - \mathbf{B} \cdot \nabla \times \nabla \times \mathbf{A})$$
$$= \int_{S} d\mathbf{S} \cdot (\mathbf{B} \times \nabla \times \mathbf{A} - \mathbf{A} \times \nabla \times \mathbf{B})$$

If S is an open surface bounded by the contour C, of which the line element is dl.

$$(33) \int_{S} d\mathbf{S} \times \nabla f = \oint_{C} d\mathbf{I} f$$

(34)
$$\int_{S} d\mathbf{S} \cdot \nabla \times \mathbf{A} = \oint_{C} d\mathbf{l} \cdot \mathbf{A}$$

(35)
$$\int_{S} (d\mathbf{S} \times \nabla) \times \mathbf{A} = \oint_{C} d\mathbf{l} \times \mathbf{A}$$

(36)
$$\int_{S} d\mathbf{S} \cdot (\nabla f \times \nabla g) = \oint_{C} f dg = -\oint_{C} g df$$

DIFFERENTIAL OPERATORS IN CURVILINEAR COORDINATES⁵

Cylindrical Coordinates

Divergence

$$\nabla \cdot \mathbf{A} = \frac{1}{r} \frac{\partial}{\partial r} (rA_r) + \frac{1}{r} \frac{\partial A_{\phi}}{\partial \phi} + \frac{\partial A_z}{\partial z}$$

Gradient

$$(\nabla f)_r = \frac{\partial f}{\partial r}; \quad (\nabla f)_\phi = \frac{1}{r} \frac{\partial f}{\partial \phi}; \quad (\nabla f)_z = \frac{\partial f}{\partial z}$$

Curl

$$(\nabla \times \mathbf{A})_r = \frac{1}{r} \frac{\partial A_z}{\partial \phi} - \frac{\partial A_\phi}{\partial z}$$

$$(\nabla \times \mathbf{A})_{\phi} = \frac{\partial A_r}{\partial z} - \frac{\partial A_z}{\partial r}$$

$$(\nabla \times \mathbf{A})_z = \frac{1}{r} \frac{\partial}{\partial r} (rA_{\phi}) - \frac{1}{r} \frac{\partial A_r}{\partial \phi}$$

Laplacian

$$\nabla^2 f = \frac{1}{r} \frac{\partial}{\partial r} \left(r \frac{\partial f}{\partial r} \right) + \frac{1}{r^2} \frac{\partial^2 f}{\partial \phi^2} + \frac{\partial^2 f}{\partial z^2}$$

Laplacian of a vector

$$(\nabla^2 \mathbf{A})_r = \nabla^2 A_r - \frac{2}{r^2} \frac{\partial A_\phi}{\partial \phi} - \frac{A_r}{r^2}$$

$$(\nabla^2 \mathbf{A})_{\phi} = \nabla^2 A_{\phi} + \frac{2}{r^2} \frac{\partial A_r}{\partial \phi} - \frac{A_{\phi}}{r^2}$$

$$(\nabla^2 \mathbf{A})_z = \nabla^2 A_z$$

Components of $(A \cdot \nabla)B$

$$(\mathbf{A} \cdot \nabla \mathbf{B})_r = A_r \frac{\partial B_r}{\partial r} + \frac{A_{\phi}}{r} \frac{\partial B_r}{\partial \phi} + A_z \frac{\partial B_r}{\partial z} - \frac{A_{\phi} B_{\phi}}{r}$$

$$(\mathbf{A} \cdot \nabla \mathbf{B})_{\phi} = A_r \frac{\partial B_{\phi}}{\partial r} + \frac{A_{\phi}}{r} \frac{\partial B_{\phi}}{\partial \phi} + A_z \frac{\partial B_{\phi}}{\partial z} + \frac{A_{\phi}B_r}{r}$$

$$(\mathbf{A} \cdot \nabla \mathbf{B})_z = A_r \frac{\partial B_z}{\partial \tau} + \frac{A_\phi}{\tau} \frac{\partial B_z}{\partial \phi} + A_z \frac{\partial B_z}{\partial z}$$

Divergence of a tensor

$$(\nabla \cdot T)_r = \frac{1}{r} \frac{\partial}{\partial r} (r T_{rr}) + \frac{1}{r} \frac{\partial T_{\phi r}}{\partial \phi} + \frac{\partial T_{rr}}{\partial z} - \frac{T_{\phi \phi}}{r}$$

$$(\nabla \cdot T)_{\phi} = \frac{1}{r} \frac{\partial}{\partial r} (r T_{r\phi}) + \frac{1}{r} \frac{\partial T_{\phi\phi}}{\partial \phi} + \frac{\partial T_{z\phi}}{\partial z} + \frac{T_{\phi r}}{r}$$

$$(\nabla \cdot T)_z = \frac{1}{r} \frac{\partial}{\partial r} (rT_{rz}) + \frac{1}{r} \frac{\partial T_{\phi z}}{\partial \phi} + \frac{\partial T_{zz}}{\partial z}$$

Spherical Coordinates

Divergence

$$\nabla \cdot \mathbf{A} = \frac{1}{r^2} \frac{\partial}{\partial r} (r^2 A_r) + \frac{1}{r \sin \theta} \frac{\partial}{\partial \theta} (\sin \theta A_{\theta}) + \frac{1}{r \sin \theta} \frac{\partial A_{\phi}}{\partial \phi}$$

Gradient

$$(\nabla f)_r = \frac{\partial f}{\partial r}; \quad (\nabla f)_\theta = \frac{1}{r} \frac{\partial f}{\partial \theta}; \quad (\nabla f)_\phi = \frac{1}{r \sin \theta} \frac{\partial f}{\partial \phi}$$

Curl

$$(\nabla \times \mathbf{A})_r = \frac{1}{r \sin \theta} \frac{\partial}{\partial \theta} (\sin \theta A_{\phi}) - \frac{1}{r \sin \theta} \frac{\partial A_{\theta}}{\partial \phi}$$

$$(\nabla \times \mathbf{A})_{\theta} = \frac{1}{r \sin \theta} \frac{\partial A_r}{\partial \phi} - \frac{1}{r} \frac{\partial}{\partial r} (r A_{\phi})$$

$$(\nabla \times \mathbf{A})_{\phi} = \frac{1}{r} \frac{\partial}{\partial r} (rA_{\theta}) - \frac{1}{r} \frac{\partial A_{r}}{\partial \theta}$$

Laplacian

$$\nabla^2 f = \frac{1}{r^2} \frac{\partial}{\partial r} \left(r^2 \frac{\partial f}{\partial r} \right) + \frac{1}{r^2 \sin \theta} \frac{\partial}{\partial \theta} \left(\sin \theta \frac{\partial f}{\partial \theta} \right) + \frac{1}{r^2 \sin^2 \theta} \frac{\partial^2 f}{\partial \phi^2}$$

Laplacian of a vector

$$(\nabla^2 \mathbf{A})_r = \nabla^2 A_r - \frac{2A_r}{r^2} - \frac{2}{r^2} \frac{\partial A_\theta}{\partial \theta} - \frac{2 \cot \theta A_\theta}{r^2} - \frac{2}{r^2 \sin \theta} \frac{\partial A_\phi}{\partial \phi}$$

$$(\nabla^2 \mathbf{A})_{\theta} = \nabla^2 A_{\theta} + \frac{2}{r^2} \frac{\partial A_r}{\partial \theta} - \frac{A_{\theta}}{r^2 \sin^2 \theta} - \frac{2 \cos \theta}{r^2 \sin^2 \theta} \frac{\partial A_{\phi}}{\partial \phi}$$

$$(\nabla^2 \mathbf{A})_{\phi} = \nabla^2 A_{\phi} - \frac{A_{\phi}}{r^2 \sin^2 \theta} + \frac{2}{r^2 \sin \theta} \frac{\partial A_r}{\partial \phi} + \frac{2 \cos \theta}{r^2 \sin^2 \theta} \frac{\partial A_{\theta}}{\partial \phi}$$

Components of $(\mathbf{A} \cdot \nabla)\mathbf{B}$

$$(\mathbf{A} \cdot \nabla \mathbf{B})_r = A_r \frac{\partial B_r}{\partial r} + \frac{A_{\theta}}{r} \frac{\partial B_r}{\partial \theta} + \frac{A_{\phi}}{r \sin \theta} \frac{\partial B_r}{\partial \phi} - \frac{A_{\theta} B_{\theta} + A_{\phi} B_{\phi}}{r}$$

$$(\mathbf{A} \cdot \nabla \mathbf{B})_{\theta} = A_r \frac{\partial B_{\theta}}{\partial r} + \frac{A_{\theta}}{r} \frac{\partial B_{\theta}}{\partial \theta} + \frac{A_{\phi}}{r \sin \theta} \frac{\partial B_{\theta}}{\partial \phi} + \frac{A_{\theta} B_r}{r} - \frac{\cot \theta A_{\phi} B_{\phi}}{r}$$

$$(\mathbf{A} \cdot \nabla \mathbf{B})_{\phi} = A_r \frac{\partial B_{\phi}}{\partial r} + \frac{A_{\theta}}{r} \frac{\partial B_{\phi}}{\partial \theta} + \frac{A_{\phi}}{r \sin \theta} \frac{\partial B_{\phi}}{\partial \phi} + \frac{A_{\phi} B_r}{r} + \frac{\cot \theta A_{\phi} B_{\theta}}{r}$$

Divergence of a tensor

$$(\nabla \cdot T)_r = \frac{1}{r^2} \frac{\partial}{\partial r} (r^2 T_{rr}) + \frac{1}{r \sin \theta} \frac{\partial}{\partial \theta} (\sin \theta T_{\theta r})$$

$$+\frac{1}{r\sin\theta}\frac{\partial T_{\phi r}}{\partial \phi}-\frac{T_{\theta\theta}+T_{\phi\phi}}{r}$$

$$(\nabla \cdot T)_{\theta} = \frac{1}{r^2} \frac{\partial}{\partial r} (r^2 T_{r\theta}) + \frac{1}{r \sin \theta} \frac{\partial}{\partial \theta} (\sin \theta T_{\theta\theta})$$

$$+\frac{1}{r\sin\theta}\frac{\partial T_{\phi\theta}}{\partial \phi}+\frac{T_{\theta r}}{r}-\frac{\cot\theta T_{\phi\phi}}{r}$$

$$(\nabla \cdot T)_{\phi} = \frac{1}{r^2} \frac{\partial}{\partial r} (r^2 T_{r\phi}) + \frac{1}{r \sin \theta} \frac{\partial}{\partial \theta} (\sin \theta T_{\theta\phi})$$

$$+\frac{1}{r\sin\theta}\frac{\partial T_{\phi\phi}}{\partial \phi}+\frac{T_{\phi r}}{r}+\frac{\cot\theta T_{\phi\theta}}{r}$$

DIMENSIONS AND UNITS

To get the value of a quantity in Gaussian units, multiply the value expressed in SI units by the conversion factor. Multiples of 3 in the conversion factors result from approximating the speed of light $c=2.9979\times 10^{10}\,\mathrm{cm/sec}\approx 3\times 10^{10}\,\mathrm{cm/sec}$.

| Dharical | S | Din | nensions | SI | Conversion | Gaussian |
|----------------------|-----------------|-----------------------|------------------------------|----------------------------|------------------------------|---------------------------------|
| Physical Quantity | Sym- bol | SI | Gaussian | Units | Factor | Units |
| Capacitance | С | $\frac{t^2q^2}{ml^2}$ | ı | farad | 9 × 10 ¹¹ | cm |
| Charge | q | q | $\frac{m^{1/2}l^{3/2}}{t}$ | coulomb | 3 × 10 ⁹ | statcoulomb |
| Charge density | ρ | $\frac{q}{l^3}$ | $\frac{m^{1/2}}{l^{3/2}t}$ | coulomb /m ³ | 3×10^3 | statcoulomb /cm³ |
| Conductance | | $\frac{tq^2}{ml^2}$ | $\frac{l}{t}$ | siemens | 9 × 10 ¹¹ | cm/sec |
| Conductivity | σ | $\frac{tq^2}{ml^3}$ | $\frac{1}{t}$ | siemens /m | 9 × 10 ⁹ | sec ⁻¹ |
| Current | I, i | $\frac{q}{t}$ | $\frac{m^{1/2}l^{3/2}}{t^2}$ | ampere | 3 × 10 ⁹ | statampere |
| Current density | J,j | $\frac{q}{l^2t}$ | $\frac{m^{1/2}}{l^{1/2}t^2}$ | ampere /m² | 3×10^5 | statampere /cin ² |
| Density | ρ | $\frac{m}{l^3}$ | $\frac{m}{l^3}$ | kg/m ³ | 10-3 | g/cm ³ |
| Displacement | D | $\frac{q}{l^2}$ | $\frac{m^{1/2}}{l^{1/2}t}$ | coulomb /m ² | $12\pi \times 10^5$ | statcoulomb /cm² |
| Electric field | E | $\frac{ml}{t^2q}$ | $\frac{m^{1/2}}{l^{1/2}t}$ | volt/m | $\frac{1}{3} \times 10^{-4}$ | statvolt/cm |
| Electro- motance | \mathcal{E} , | $\frac{ml^2}{t^2q}$ | $\frac{m^{1/2}l^{1/2}}{t}$ | volt | $\frac{1}{3}\times 10^{-2}$ | statvolt |
| Energy | U, W | $\frac{ml^2}{t^2}$ | $\frac{ml^2}{t^2}$ | joule | 107 | erg |
| Energy density | w, e | $\frac{m}{lt^2}$ | $\frac{m}{lt^2}$ | joule/m³ | 10 | erg/cm ³ |

| Division | C | Din | nensions | SI | C | |
|----------------------|-------------|---------------------|------------------------------|----------------------|-------------------------------|-----------------------------|
| Physical Quantity | Sym- bol | SI | Gaussian | Units | Conversion Factor | Gaussian Units |
| Force | F | $\frac{ml}{t^2}$ | $\frac{ml}{t^2}$ | newton | 10 ⁵ | dyne |
| Frequency | f, u | $\frac{1}{t}$ | $\frac{1}{t}$ | hertz | 1 | hertz |
| Impedance | Z | $\frac{ml^2}{tq^2}$ | $\frac{t}{l}$ | ohm | $\frac{1}{9}\times10^{-11}$ | sec/cm |
| Inductance | L | $\frac{ml^2}{q^2}$ | $\frac{t^2}{l}$ | henry | $\frac{1}{9} \times 10^{-11}$ | sec²/cm |
| Length | I | ı | ı | meter (m) | 10 ² | centimeter (cm) |
| Magnetic intensity | н | $\frac{q}{lt}$ | $\frac{m^{1/2}}{l^{1/2}t}$ | ampere- turn/m | $4\pi \times 10^{-3}$ | oersted |
| Magnetic flux | Φ | $\frac{ml^2}{tq}$ | $\frac{m^{1/2}l^{3/2}}{t}$ | weber | 10 ⁸ | maxwell |
| Magnetic induction | в | $\frac{m}{tq}$ | $\frac{m^{1/2}}{l^{1/2}t}$ | tesla | 104 | gauss |
| Magnetic moment | m, μ | $\frac{l^2q}{t}$ | $\frac{m^{1/2}l^{5/2}}{t}$ | ampere-m² | 10 ³ | oersted- cm ³ |
| Magnetization | М | $\frac{q}{lt}$ | $\frac{m^{1/2}}{l^{1/2}t}$ | ampere- turn/m | 10-3 | oersted |
| Magneto- motance | M, Mmf | $\frac{q}{t}$ | $\frac{m^{1/2}l^{1/2}}{t^2}$ | ampere- turn | $\frac{4\pi}{10}$ | gilbert |
| Mass | m, M | m | m | kilogram (kg) | 10 ³ | gram (g) |
| Momentum | p, P | $\frac{ml}{t}$ | $\frac{ml}{t}$ | kg-m/s | 10 ⁵ | g-cm/sec |
| Momentum density | | $\frac{m}{l^2t}$ | $\frac{m}{l^2t}$ | kg/m ² -s | 10-1 | g/cm ² -sec |
| Permeability | μ | $\frac{ml}{q^2}$ | 1 | henry/m | $\frac{1}{4\pi} \times 10^7$ | |

| Physical | Sym- | Din | nensions | SI | Conversion | Gaussian | |
|---------------------------|-------------|-----------------------|---|------------------------|-----------------------------|--------------------------|--|
| Quantity | bol | SI | Gaussian | Units | Factor | Units | |
| Permittivity | € | $\frac{t^2q^2}{ml^3}$ | 1 | farad/m | $36\pi \times 10^9$ | | |
| Polarization | P | $\frac{q}{l^2}$ | $\frac{m^{1/2}}{l^{1/2}t}$ | coulomb/m ² | 3×10^5 | statcoulomb /cm² | |
| Potential | V,ϕ | $\frac{ml^2}{t^2q}$ | $\frac{m^{1/2}l^{1/2}}{t}$ | volt | $\frac{1}{3}\times10^{-2}$ | statvolt | |
| Power | P | $\frac{ml^2}{t^3}$ | $\frac{ml^2}{t^3}$ | watt | 107 | erg/sec | |
| Power density | | $\frac{m}{lt^3}$ | $\frac{m}{lt^3}$ | watt/m ³ | 10 | erg/cm ³ -sec | |
| Pressure | p, P | $\frac{m}{lt^2}$ | $\frac{m}{lt^2}$ | pascal | 10 | dyne/cm² | |
| Reluctance | R | $\frac{q^2}{ml^2}$ | $\frac{1}{l}$ | ampere-turn /weber | $4\pi\times10^{-9}$ | cm ⁻¹ | |
| Resistance | R | $\frac{ml^2}{tq^2}$ | $\frac{t}{l}$ | ohm | $\frac{1}{9}\times10^{-11}$ | sec/cm | |
| Resistivity | η, ρ | $\frac{ml^3}{tq^2}$ | t | ohm-m | $\frac{1}{9}\times10^{-9}$ | sec | |
| Thermal con- ductivity | κ, k | $\frac{ml}{t^3}$ | $\frac{ml}{t^3}$ | watt/m- deg (K) | 10 ⁵ | erg/cm-sec- deg (K) | |
| Time | t | t | t | second (s) | 1 | second (sec) | |
| Vector potential | A | $\frac{ml}{tq}$ | $\frac{m^{1/2}l^{1/2}}{t}$ | weber/m | 10 ⁶ | gauss-cm | |
| Velocity | v | $\frac{l}{t}$ | $\left\{ \begin{array}{l} rac{l}{t} \end{array} \right.$ | m/s | 10 ² | cm/sec | |
| Viscosity | η, μ | $\frac{m}{lt}$ | $\frac{m}{lt}$ | kg/ni-s | 10 | poise | |
| Vorticity | ζ | $\frac{1}{t}$ | $\frac{1}{t}$ | s ⁻¹ | 1 | sec ⁻¹ | |
| Work | W | $\frac{ml^2}{t^2}$ | $\frac{ml^2}{t^2}$ | joule | 107 | erg | |

INTERNATIONAL SYSTEM (SI) NOMENCLATURE⁶

| Physical Quantity | Name of Unit | Symbol for Unit | Physical Quantity | Name of Unit | Symbol for Unit |
|------------------------|-----------------|--------------------|----------------------------|-----------------|--------------------|
| *length | meter | m | electric | volt | V |
| *mass | kilogram | kg | potential | | |
| *time | second | s | electric resistance | ohm | Ω |
| *current | ampere | A | electric | siemens | s |
| *temperature | kelvin | K | conductance | | _ |
| *amount of | mole | mol | electric capacitance | farad | F |
| substance | | , | magnetic flux | weber | Wb |
| *luminous intensity | candela | cd | magnetic | henry | н |
| †plane angle | radian | rad | inductance | | |
| tsolid angle | steradian | sr | magnetic intensity | tesla | T |
| frequency | hertz | Hz | luminous flux | lumen | $_{ m lm}$ |
| energy | joule | J | illuminance | lux | lx |
| force | newton | N | activity (of a radioactive | becquerel | Bq |
| pressure | pascal | Pa | source) | | |
| power | watt | W | absorbed dose | gray | Gy |
| electric charge | coulomb | C | (of ionizing radiation) | | |

METRIC PREFIXES

| Multiple | Prefix | Symbol | Multiple | Prefix | Symbol |
|----------|--------|--------|------------------|--------|--------|
| 10-1 | deci | d | 10 | deca | da |
| 10-2 | centi | С | 10 ² | hecto | h |
| 10-3 | milli | m | 10 ³ | kilo | k |
| 10-6 | micro | μ | 10 ⁶ | mega | M |
| 10-9 | nano | n | 10 ⁹ | giga | G |
| 10-12 | pico | P | 1012 | tera | T |
| 10-15 | femto | f | 10 ¹⁵ | peta | P |
| 10-18 | atto | a | 10 ¹⁸ | exa | E |

PHYSICAL CONSTANTS (SI)⁷

| Physical Quantity | Symbol | Value | Units |
|------------------------------|---|--|--------------------|
| Boltzmann constant | k | 1.3807×10^{-23} | J K-1 |
| Elementary charge | e | 1.6022×10^{-19} | C |
| Electron mass | m_{e} | 9.1095×10^{-31} | kg |
| Proton mass | m_p | 1.6726×10^{-27} | kg |
| Gravitational constant | G | 6.6720×10^{-11} | $m^3s^{-2}kg^{-1}$ |
| Planck constant | h | 6.6262×10^{-34} | Js |
| 1 | $\hbar = h/2\pi$ | 1.0546×10^{-34} | Js |
| Speed of light in vacuum | c | 2.9979×10^8 | $m s^{-1}$ |
| Permittivity of free space | €0 | 8.8542×10^{-12} | F m ⁻¹ |
| Permeability of free space | μο | $4\pi\times10^{-7}$ | H m ⁻¹ |
| Proton/electron mass ratio | m_p/m_e | 1.8362×10^3 | |
| Electron charge/mass ratio | e/m _e | 1.7588×10^{11} | C kg ⁻¹ |
| Rydberg constant | $R_{\infty} = \frac{me^4}{8\epsilon_0^2 ch^3}$ | 1.0974×10^7 | m ⁻¹ |
| Bohr radius | $a_0 = \epsilon_0 h^2 / \pi m e^2$ | 5.2918×10^{-11} | m |
| Atomic cross section | πa_0^2 | 8.7974×10^{-21} | m ² |
| Classical electron radius | $r_e = e^2/4\pi\epsilon_0 mc^2$ | 2.8179×10^{-15} | m |
| Thomson cross section | $(8\pi/3)r_e^2$ | 6.6524×10^{-29} | m² |
| Compton wavelength of | h/mec | 2.4263×10^{-12} | m |
| electron | ħ/mec | 3.8616×10^{-13} | ın |
| Fine-structure constant | $\begin{array}{c} \alpha = e^2/2\epsilon_0 hc \\ \alpha^{-1} \end{array}$ | $\begin{array}{c c} 7.2974 \times 10^{-3} \\ 137.04 \end{array}$ | |
| First radiation constant | $c_1 = 2\pi h c^2$ | 3.7418×10^{-2} | W m ² |
| Second radiation constant | $c_2 = hc/k$ | 1.4388×10^{-2} | m K |
| Stefan-Boltzmann constant | σ | 5.6703×10^{-8} | W m - 2 K - 4 |

| Physical Quantity | Symbol | Value | Units |
|--|-------------------------|--------------------------|------------------------------------|
| Wavelength associated with 1 eV | $\lambda_0 = hc/e$ | 1.2399×10^{-6} | m |
| Frequency associated with 1 eV | $\nu_0 = e/h$ | 2.4180×10^{14} | Hz |
| Wave number associated with 1 eV | $k_0 = e/hc$ | 8.0655×10^5 | m ⁻¹ |
| Energy associated with 1 eV | $h u_0$ | 1.6022×10^{-19} | J |
| Energy associated with 1 m ⁻¹ | h c | 1.9865×10^{-25} | J |
| Energy associated with 1 Rydberg | $me^3/8\epsilon_0^2h^2$ | 13.606 | eV |
| Energy associated with 1 Kelvin | k/e | 8.6173×10^{-5} | eV |
| Temperature associated with 1 eV | e/k | 1.1605×10^4 | K |
| Avogadro number | N_A | 6.0220×10^{23} | mol-1 |
| Faraday constant | $F = N_A e$ | 9.6485×10^4 | C mol-1 |
| Gas constant | $R = N_A k$ | 8.3144 | JK ⁻¹ mol ⁻¹ |
| Loschmidt's number (no. density at STP) | n_0 | 2.6868×10^{25} | m ⁻³ |
| Atomic mass unit | m_u | 1.6606×10^{-27} | kg |
| Standard temperature | T_0 | 273.16 | K |
| Atmospheric pressure | $p_0 = n_0 k T_0$ | 1.0133×10^{5} | Pa |
| Pressure of 1 mm Hg (1 torr) | | 1.3332×10^2 | Pa |
| Molar volume at STP | $V_0 = RT_0/p_0$ | 2.2415×10^{-2} | m ³ |
| Molar weight of air | Mair | 2.8971×10^{-2} | kg |
| calorie (cal) | | 4.1868 | J |
| Gravitational acceleration | g | 9.8067 | m s ⁻² |

PHYSICAL CONSTANTS (cgs)

| Physical Quantity | Symbol | Value | Units |
|---------------------------------|--|----------------------------------|---|
| Boltzmann constant | k | 1.3807×10^{-16} | erg/deg(K) |
| Elementary charge | e | 4.8032×10^{-10} | statcoulomb (statcoul) |
| Electron mass | m_e | 9.1095×10^{-28} | g |
| Proton mass | m_p | 1.6726×10^{-24} | g |
| Gravitational constant | G | 6.6720×10^{-8} | dyne-cm ² /g ² |
| Planck constant | h | 6.6262×10^{-27} | erg-sec |
| 1 | $\hbar = h/2\pi$ | 1.0546×10^{-27} | erg-sec |
| Speed of light in vacuum | c | 2.9979×10^{10} | cm/sec |
| Proton/electron mass ratio | m_p/m_e | 1.8362×10^3 | |
| Electron charge/mass ratio | e/m _e | 5.2728×10^{17} | statcoul/g |
| Rydberg constant | $R_{\infty} = \frac{2\pi^2 m e^4}{ch^3}$ | 1.0974×10^{5} | cm ⁻¹ |
| Bohr radius | $a_0 = \hbar^2/me^2$ | 5.2918×10^{-9} | cm |
| Atomic cross section | πa_0^2 | 8.7974×10^{-17} | cm ² |
| Classical electron radius | $r_e = e^2/mc^2$ | 2.8179×10^{-13} | cm |
| Thomson cross section | $(8\pi/3)r_e^2$ | 6.6524×10^{-25} | cm ² |
| Compton wavelength of | h/mec | 2.4263×10^{-10} | cm |
| electron | ħ/mec | 3.8616×10^{-11} | cm |
| Fine-structure constant | $\alpha = e^2/\hbar c$ α^{-1} | 7.2974×10^{-3} 137.04 | |
| First radiation constant | $c_1 = 2\pi h c^2$ | 3.7418×10^{-5} | erg-cm ² /sec |
| Second radiation constant | $c_2 = hc/k$ | 1.4388 | cm-deg(K) |
| Stefan-Boltzmann constant | σ | 5.6703×10^{-5} | erg/cm ² - sec-deg ⁴ |
| Wavelength associated with 1 eV | λ ₀ | 1.2399 × 10 ⁻⁴ | cm |

| Physical Quantity | Symbol | Value | Units |
|---|-------------------|--------------------------|----------------------|
| Frequency associated with 1 eV | ν ₀ | 2.4180×10^{14} | Нz |
| Wave number associated with 1 eV | k_0 | 8.0655×10^3 | cm ⁻¹ |
| Energy associated with 1 eV | | 1.6022×10^{-12} | erg |
| Energy associated with 1 cm ⁻¹ | | 1.9865×10^{-16} | erg |
| Energy associated with 1 Rydberg | | 13.606 | eV |
| Energy associated with 1 deg Kelvin | | 8.6173×10^{-5} | eV |
| Temperature associated with 1 eV | | 1.1605 × 10 ⁴ | deg(K) |
| Avogadro number | N_A | 6.0220×10^{23} | mol ⁻¹ |
| Faraday constant | $F = N_A e$ | 2.8925×10^{14} | statcoul/mol |
| Gas constant | $R = N_A k$ | 8.3144×10^7 | erg/deg-mol |
| Loschmidt's number (no. density at STP) | n_0 | 2.6868×10^{19} | cm ⁻³ |
| Atomic mass unit | $m_{\mathbf{u}}$ | 1.6606×10^{-24} | g |
| Standard temperature | T_0 | 273.16 | deg(K) |
| Atmospheric pressure | $p_0 = n_0 k T_0$ | 1.0133×10^{6} | dyne/cm ² |
| Pressure of 1 mm Hg (1 torr) | | 1.3332×10^3 | dyne/cm ² |
| Molar volume at STP | $V_0 = RT_0/p_0$ | 2.2415×10^4 | cm ³ |
| Molar weight of air | Mair | 28.971 | g |
| calorie (cal) | 1 | 4.1868×10^{7} | erg |
| Gravitational acceleration | g | 980.67 | cm/sec ² |

FORMULA CONVERSION8

Here $\alpha=10^2$ cm m⁻¹, $\beta=10^7$ erg J⁻¹, $\epsilon_0=8.8542\times 10^{-12}$ F m⁻¹, $\mu_0=4\pi\times 10^{-7}$ H m⁻¹, $c=(\epsilon_0\mu_0)^{-1/2}=2.9979\times 10^8$ m s⁻¹, and $\hbar=1.0546\times 10^{-34}$ J s. To derive a dimensionally correct SI formula from one expressed in Gaussian units, substitute for each quantity according to $\bar{Q}=\bar{k}Q$, where \bar{k} is the coefficient in the second column of the table corresponding to Q (overbars denote variables expressed in Gaussian units). Thus, the formula $\bar{a}_0=\bar{h}^2/\bar{m}\bar{e}^2$ for the Bohr radius becomes $\alpha a_0=(\hbar\beta)^2/[(m\beta/\alpha^2)(e^2\alpha\beta/4\pi\epsilon_0)]$, or $a_0=\epsilon_0h^2/\pi me^2$. To go from SI to natural units in which $\hbar=c=1$ (distinguished by a circumflex), use $Q=\hat{k}^{-1}\hat{Q}$, where \hat{k} is the coefficient corresponding to Q in the third column. Thus $\hat{a}_0=4\pi\epsilon_0\hbar^2/[(\hat{m}\hbar/c)(\hat{e}^2\epsilon_0\hbar c)]=4\pi/\hat{m}\hat{e}^2$. (In transforming from SI units, do not substitute for ϵ_0 , μ_0 , or ϵ .)

MAXWELL'S EQUATIONS

| Name or Description | SI | Gaussian |
|---------------------------------|--|---|
| Faraday's law | $\nabla \times \mathbf{E} = -\frac{\partial \mathbf{B}}{\partial t}$ | $\nabla \times \mathbf{E} = -\frac{1}{c} \frac{\partial \mathbf{B}}{\partial t}$ |
| Ampere's law | $\nabla \times \mathbf{H} = \frac{\partial \mathbf{D}}{\partial t} + \mathbf{J}$ | $\nabla \times \mathbf{H} = \frac{1}{c} \frac{\partial \mathbf{D}}{\partial t} + \frac{4\pi}{c} \mathbf{J}$ |
| Poisson equation | $\nabla \cdot \mathbf{D} = \rho$ | $\nabla \cdot \mathbf{D} = 4\pi \rho$ |
| [Absence of magnetic monopoles] | $\nabla \cdot \mathbf{B} = 0$ | $\nabla \cdot \mathbf{B} = 0$ |
| Lorentz force on charge q | $q(E + v \times B)$ | $q\left(\mathbf{E} + \frac{1}{c}\mathbf{v} \times \mathbf{B}\right)$ |
| Constitutive relations | $D = \epsilon E$ $B = \mu H$ | $D = \epsilon E$ $B = \mu H$ |

In a plasma, $\mu \approx \mu_0 = 4\pi \times 10^{-7} \, \mathrm{H \, m^{-1}}$ (Gaussian units: $\mu \approx 1$). The permittivity satisfies $\epsilon \approx \epsilon_0 = 8.8542 \times 10^{-12} \, \mathrm{F \, m^{-1}}$ (Gaussian: $\epsilon \approx 1$) provided that all charge is regarded as free. Using the drift approximation $\mathbf{v}_{\perp} = \mathbf{E} \times \mathbf{B}/B^2$ to calculate polarization charge density gives rise to a dielectric constant $K \equiv \epsilon/\epsilon_0 = 1 + 36\pi \times 10^9 \, \rho/B^2$ (SI) $= 1 + 4\pi \rho c^2/B^2$ (Gaussian), where ρ is the mass density.

The electromagnetic energy in volume V is given by

$$W = \frac{1}{2} \int_{V} dV (\mathbf{H} \cdot \mathbf{B} + \mathbf{E} \cdot \mathbf{D})$$
 (SI)
$$= \frac{1}{8\pi} \int_{V} dV (\mathbf{H} \cdot \mathbf{B} + \mathbf{E} \cdot \mathbf{D})$$
 (Gaussian).

Poynting's theorem is

$$\frac{\partial W}{\partial t} + \int_{S} \mathbf{N} \cdot d\mathbf{S} = -\int_{V} dV \mathbf{J} \cdot \mathbf{E},$$

where S is the closed surface bounding V and the Poynting vector (energy flux across S) is given by $N = E \times H$ (SI) or $N = cE \times H/4\pi$ (Gaussian).

ELECTRICITY AND MAGNETISM

In the following, $\epsilon =$ dielectric permittivity, $\mu =$ permeability of conductor, $\mu' =$ permeability of surrounding medium, $\sigma =$ conductivity, $f = \omega/2\pi =$ radiation frequency, $\kappa_m = \mu/\mu_0$ and $\kappa_e = \epsilon/\epsilon_0$. Where subscripts are used, '1' denotes a conducting medium and '2' a propagating (lossless dielectric) medium. All units are SI unless otherwise specified.

| , | • |
|---|---|
| Permittivity of free space | $\epsilon_0 = 8.8542 \times 10^{-12} \mathrm{F m^{-1}}$ |
| Permeability of free space | $\mu_0 = 4\pi \times 10^{-7} \mathrm{H m^{-1}}$ = 1.2566 × 10 ⁻⁶ H m ⁻¹ |
| Resistance of free space | $R_0 = (\mu_0/\epsilon_0)^{1/2} = 376.73\Omega$ |
| Capacity of parallel plates of area A, separated by distance d | $C = \epsilon A/d$ |
| Capacity of concentric cylinders of length l, radii a, b | $C = 2\pi \epsilon l \ln(b/a)$ |
| Capacity of concentric spheres of radii a, b | $C = 4\pi \epsilon ab/(b-a)$ |
| Self-inductance of wire of length l, carrying uniform current | $L = \mu l$ |
| Mutual inductance of parallel wires of length l , radius a , separated by distance d | $L = (\mu' l/4\pi) [1 + 4 \ln(d/a)]$ |
| Inductance of circular loop of radius b, made of wire of radius a, carrying uniform current | $L = b \left\{ \mu' \left[\ln(8b/a) - 2 \right] + \mu/4 \right\}$ |
| Relaxation time in a lossy medium | $\tau = \epsilon/\sigma$ |
| Skin depth in a lossy medium | $\delta = (2/\omega\mu\sigma)^{1/2} = (\pi f\mu\sigma)^{-1/2}$ |
| Wave impedance in a lossy medium | $Z = \left[\mu/(\epsilon + i\sigma/\omega)\right]^{1/2}$ |
| Transmission coefficient at conducting surface (good only for $T \ll 1$) | $T = 4.22 \times 10^{-4} (f \kappa_{m1} \kappa_{e2} / \sigma)^{1/2}$ |
| Field at distance r from straight wire carrying current I (amperes) | $B_{\theta} = \mu I/2\pi r \text{ tesla}$ = 0.2I/r gauss (r in cm) |
| Field at distance z along axis from circular loop of radius a | $B_z = \mu a^2 I / [2(a^2 + z^2)^{3/2}]$ |

carrying current I

ELECTROMAGNETIC FREQUENCY/ WAVELENGTH BANDS¹⁰

| | Frequency Range | | Wavelength Range | |
|---------------|-----------------|---------|------------------|--------|
| Designation | Lower | Upper | Lower | Upper |
| ULF* | | 10 Hz | 3 Mm | |
| ELF* | 10 Hz | 3 kHz | 100 km | 3 Min |
| VLF | 3 kHz | 30 kHz | 10 km | 100 km |
| LF | 30 kHz | 300 kHz | 1 km | 10 km |
| MF | 300 kHz | 3 MHz | 100 m | 1 km |
| HF | 3 MHz | 30 MHz | 10 m | 100 m |
| VHF | 30 MHz | 300 MHz | 1 m | 10 m |
| UHF | 300 MHz | 3 GHz | 10 cm | 1 m |
| SHF† | 3 GHz | 30 GHz | 1 cm | 10 cm |
| S | 2.6 | 3.95 | 7.6 | 11.5 |
| G | 3.95 | 5.85 | 5.1 | 7.6 |
| J | 5.3 | 8.2 | 3.7 | 5.7 |
| H | 7.05 | 10.0 | 3.0 | 4.25 |
| x | 8.2 | 12.4 | 2.4 | 3.7 |
| M | 10.0 | 15.0 | 2.0 | 3.0 |
| P | 12.4 | 18.0 | 1.67 | 2.4 |
| K | 18.0 | 26.5 | 1.1 | 1.67 |
| R | 26.5 | 40.0 | 0.75 | 1.1 |
| EHF | 30 GHz | 300 GHz | 1 mm | 1 cm |
| Submillimeter | 300 GHz | 3 THz | 100 µm | 1 mm |
| Infrared | 3 THz | 430 THz | 700 nm | 100 µm |
| Visible | 430 THz | 750 THz | 400 nm | 700 nm |
| Ultraviolet | 750 THz | 30 PHz | 10 nm | 400 nm |
| X Ray | 30 PHz | 3 EHz | 100 pm | 10 nm |
| Gamma Ray | 3 EHz | | | 100 pm |

Note: In spectroscopy the angstrom (Å) is sometimes used $(1 \text{ Å} = 10^{-8} \text{ cm} = 0.1 \text{ nm})$.

^{*}The boundary between ULF and ELF is variously defined.

[†]The SHF (microwave) band is further subdivided approximately as shown.11

AC CIRCUITS

For a resistance R, inductance L, and capacitance C in series with a voltage source $V=V_0\exp(i\omega t)$ (here $i=\sqrt{-1}$), the current is given by I=dq/dt, where q satisfies

$$L\frac{d^2q}{dt^2} + R\frac{dq}{dt} + \frac{q}{C} = V.$$

Solutions are $q(t)=q_s+q_t$, $I(t)=I_s+I_t$, where the steady state is $I_s=i\omega q_s=V/Z$ in terms of the impedance $Z=R+i(\omega L-1/\omega C)$ and $I_t=dq_t/dt$. For initial conditions $q(0)\equiv q_0=\bar{q}_0+q_s$, $I(0)\equiv I_0$, the transients can be of three types, depending on $\Delta=R^2-4L/C$:

(a) Overdamped, $\Delta > 0$

$$\begin{split} q_{t} &= \frac{I_{0} + \gamma_{+} \bar{q}_{0}}{\gamma_{+} - \gamma_{-}} \exp(-\gamma_{-}t) - \frac{I_{0} + \gamma_{-} \bar{q}_{0}}{\gamma_{+} - \gamma_{-}} \exp(-\gamma_{+}t), \\ I_{t} &= \frac{\gamma_{+} (I_{0} + \gamma_{-} \bar{q}_{0})}{\gamma_{+} - \gamma_{-}} \exp(-\gamma_{+}t) - \frac{\gamma_{-} (I_{0} + \gamma_{+} \bar{q}_{0})}{\gamma_{+} - \gamma_{-}} \exp(-\gamma_{-}t), \end{split}$$

where $\gamma_{\pm} = (R \pm \Delta^{1/2})/2L$;

(b) Critically damped, $\Delta = 0$

$$q_t = \{\bar{q}_0 + (I_0 + \gamma_R \bar{q}_0)t\} \exp(-\gamma_R t),$$

$$I_t = [I_0 - (I_0 + \gamma_R \bar{q}_0)\gamma_R t] \exp(-\gamma_R t),$$

where $\gamma_R = R/2L$;

(c) Underdamped, $\Delta < 0$

$$\begin{split} q_t &= \left[\frac{\gamma_R \bar{q}_0 + I_0}{\omega_1} \sin \omega_1 t + \bar{q}_0 \cos \omega_1 t \right] \exp(-\gamma_R t), \\ I_t &= \left[I_0 \cos \omega_1 t - \frac{(\omega_1^2 + \gamma_R^2) \bar{q}_0 + \gamma_R I_0}{\omega_1} \sin(\omega_1 t) \right] \exp(-\gamma_R t). \end{split}$$

where $\omega_1 = \omega_0 (1 - R^2 C/4L)^{1/2}$ and $\omega_0 = (LC)^{-1/2}$ is the resonant frequency. At $\omega = \omega_0$, Z = R. The quality of the circuit is $Q = \omega_0 L/R$. Instability results when L, R, C are not all of the same sign.

DIMENSIONLESS NUMBERS OF FLUID MECHANICS¹²

| Name(s) | Symbol | Definition | Significance |
|--------------------|--------|--|--|
| Alfvén, Kármán | Al, Ka | V_A/V | *(Magnetic force/ inertial force) ^{1/2} |
| Bond | Bd | $(\rho'-\rho)L^2g/\Sigma$ | Gravitational force/ surface tension |
| Boussinesq | В | $V/(2gR)^{1/2}$ | (Inertial force/ gravitational force) ^{1/2} |
| Brinkman | Br | $\mu V^2/k\Delta T$ | Viscous heat/conducted heat |
| Capillary | Ср | $\mu V/\Sigma$ | Viscous force/surface tension |
| Carnot | Ca | $(T_2-T_1)/T_2$ | Theoretical Carnot cycle efficiency |
| Cauchy, Hooke | Cy, Hk | $\rho V^2/\Gamma = M^2$ | Inertial force/ compressibility force |
| Clausius | Cl | $LV^3 ho/k\Delta T$ | Kinetic energy flow rate/heat conduction rate |
| Cowling | С | $(V_A/V)^2 = \mathrm{Al}^2$ | Magnetic force/inertial force |
| Crispation | Cr | $\mu\kappa/\Sigma L$ | Effect of diffusion/effect of surface tension |
| Dean | D | $D^{3/2}V/\nu(2r)^{1/2}$ | Transverse flow due to curvature/longitudinal flow |
| [Drag coefficient] | C_D | $\frac{(\rho'-\rho)Lg/}{\rho'V^2}$ | Drag force/inertial force |
| Eckert | E | $V^2/c_p\Delta T$ | Kinetic energy/change in thermal energy |
| Ekman | Ek | $(\nu/2\Omega L^2)^{1/2} = (\text{Ro/Re})^{1/2}$ | (Viscous force/Coriolis force) ^{1/2} |
| Euler | Eu | $\Delta p/\rho V^2$ | Pressure drop due to friction/ dynamic pressure |
| Froude | Fr | $V/(gL)^{1/2} \ V/NL$ | †(Inertial force/gravitational or buoyancy force) ^{1/2} |
| Gay-Lussac | Ga | $1/eta\Delta T$ | Inverse of relative change in volume during heating |
| Grashof | Gr | $gL^3\beta\Delta T/\nu^2$ | Buoyancy force/viscous force |
| [Hall coefficient] | C_H | λ/r_L | Gyrofrequency/ collision frequency |

^{*(†)} Also defined as the inverse (square) of the quantity shown.

| Name(s) | Symbol | Definition | Significance |
|----------------------|--------|---|---|
| Hartmann | Н | $\frac{BL/(\mu\eta)^{1/2}}{(\operatorname{Rm}\operatorname{Re}C)^{1/2}}$ | Magnetic force/ dissipative force |
| Knudsen | Kn | λ/L | Hydrodynamic time/ collision time |
| Lorentz | Lo | V/c | Magnitude of relativistic effects |
| Lundquist | Lu | $\mu_0 L V_A / \eta = $ Al Rm | J × B force/resistive magnetic diffusion force |
| Mach | M | V/C_S | Magnitude of compressibility effects |
| Magnetic Mach | Mm | $V/V_A = Al^{-1}$ | (Inertial force/magnetic force) ^{1/2} |
| Magnetic Reynolds | Rm | $\mu_0 LV/\eta$ | Flow velocity/magnetic diffusion velocity |
| Newton | Nt | $F/\rho L^2 V^2$ | Imposed force/inertial force |
| Nusselt | N | $\alpha L/k$ | Total heat transfer/thermal conduction |
| Péclet | Pe | LV/κ | Heat convection/heat conduction |
| Poisseuille | Po | $D^2 \Delta p / \mu LV$ | Pressure force/viscous force |
| Prandtl, Schmidt | Pr, Sc | ν/κ | Momentum diffusion/ heat diffusion |
| Rayleigh | Ra | $gH^3\beta\Delta T/\nu\kappa$ | Buoyancy force/diffusion force |
| Reynolds | Re | LV/ u | Inertial force/viscous force |
| Richardson | Ri | $(NH/\Delta V)^2$ | Buoyancy effects/ vertical shear effects |
| Rossby | Ro | $V/2\Omega L \sin \Lambda$ | Inertial force/Coriolis force |
| Stanton | St | $\alpha/\rho c_p V$ | Thermal conduction loss/ heat capacity |
| Stefan | Sf | $\sigma LT^3/k$ | Radiated heat/conducted heat |
| Stokes | S | $\nu/L^2 f$ | Viscous damping rate/ vibration frequency |
| Strouhal | Sr | fL/V | Vibration speed/flow velocity |
| Taylor | Та | $\begin{pmatrix} (2\Omega L^2/\nu)^2 \\ R^{1/2} (\Delta R)^{3/2} \\ \cdot (\Omega/\nu) \end{pmatrix}$ | Centrifugal force/viscous force (Centrifugal force/ viscous force) ^{1/2} |
| Thring. Boltzmann | Th. Bo | $\rho c_p V / \epsilon \sigma T^3$ | Convective heat transport/ radiative heat transport |
| Weber | w | $\rho LV^2/\Sigma$ | Inertial force/surface tension |

Nomenclature:

| В | Magnetic induction |
|--|--|
| C_s, c | Speeds of sound, light |
| c_p | Specific heat at constant pressure (units $m^2 s^{-2} K^{-1}$) |
| D = 2R | Pipe diameter |
| F | Imposed force |
| f | Vibration frequency |
| g | Gravitational acceleration |
| H, L | Vertical, horizontal length scales |
| $k = \rho c_p \kappa$ | Thermal conductivity (units $kg m^{-1} s^{-2}$) |
| $N = (g/H)^{1/2}$ | Brunt-Väisälä frequency |
| R | Radius of pipe or channel |
| r | Radius of curvature of pipe or channel |
| r_L | Larmor radius |
| T | Temperature |
| V | Characteristic flow velocity |
| $V_A = B/(\mu_0 \rho)^{1/2}$ | Alfvén speed |
| α | Newton's-law heat coefficient, $k \frac{\partial T}{\partial x} = \alpha \Delta T$ |
| β | Volumetric expansion coefficient, $dV/V = \beta dT$ |
| Γ | Bulk modulus (units kg m ⁻¹ s ⁻²) |
| $\Delta R, \Delta V, \Delta p, \Delta T$ | Imposed difference in two radii, velocities, pressures, or temperatures |
| ϵ | Surface emissivity |
| η | Electrical resistivity |
| κ | Thermal diffusivity (units m ² s ⁻¹) |
| Λ | Latitude of point on earth's surface |
| λ | Collisional mean free path |
| $\mu = \rho \nu$ | Bulk viscosity |
| μ_0 | Permeability of free space |
| ν | Kinematic viscosity (units m ² s ⁻¹) |
| ρ | Mass density of fluid medium |
| ho' | Mass density of bubble, droplet, or moving object |
| Σ | Surface tension (units kg s ⁻²) |
| σ | Stefan-Boltzmann constant |
| Ω | Solid-body rotational angular velocity |
| | |

SHOCKS

At a shock front propagating in a magnetized fluid at an angle θ with respect to the magnetic induction B, the jump conditions are 13.14

(1)
$$\rho U = \bar{\rho} \bar{U} \equiv q$$
;

(2)
$$\rho U^2 + p + B_{\perp}^2/2\mu = \bar{\rho}\bar{U}^2 + \bar{p} + \bar{B}_{\perp}^2/2\mu$$
;

(3)
$$\rho UV - B_{\parallel}B_{\perp}/\mu = \bar{\rho}\bar{U}\bar{V} - \bar{B}_{\parallel}\bar{B}_{\perp}/\mu$$
;

(4)
$$B_{||} = \bar{B}_{||};$$

(5)
$$UB_{\perp} - VB_{||} = \bar{U}\bar{B}_{\perp} - \bar{V}\bar{B}_{||};$$

(6)
$$\frac{1}{2}(U^2 + V^2) + w + (UB_{\perp}^2 - VB_{\parallel}B_{\perp})/\mu\rho U$$

= $\frac{1}{2}(\bar{U}^2 + \bar{V}^2) + \bar{w} + (\bar{U}\bar{B}_{\perp}^2 - \bar{V}\bar{B}_{\parallel}\bar{B}_{\perp})/\mu\bar{\rho}\bar{U}$.

Here U and V are components of the fluid velocity normal and tangential to the front in the shock frame; $\rho = 1/v$ is the mass density; p is the pressure; $B_{\perp} = B \sin \theta$, $B_{\parallel} = B \cos \theta$; μ is the magnetic permeability ($\mu = 4\pi$ in cgs units); and the specific enthalpy is w = e + pv, where the specific internal energy e satisfies de = Tds - pdv in terms of the temperature T and the specific entropy s. Quantities in the region behind (downstream from) the front are distinguished by a bar. If B = 0, then 15

(7)
$$U - \bar{U} = [(\bar{p} - p)(\upsilon - \bar{\upsilon})]^{1/2};$$

(8)
$$(\bar{p}-p)(v-\bar{v})^{-1}=q^2;$$

(9)
$$\bar{w} - w = \frac{1}{2}(\bar{p} - p)(v + \bar{v});$$

(10)
$$\bar{e} - e = \frac{1}{2}(\bar{p} + p)(\upsilon - \bar{\upsilon}).$$

In what follows we assume that the fluid is a perfect gas with adiabatic index $\gamma = 1 + 2/n$, where n is the number of degrees of freedom. Then $p = \rho RT/m$, where R is the universal gas constant and m is the molar weight; the sound speed is given by $C_s^2 = (\partial p/\partial \rho)_s = \gamma pv$; and $w = \gamma e = \gamma pv/(\gamma + 1)$. For a general oblique shock in a perfect gas the quantity $X = r^{-1}(U/V_A)^2$ satisfies 14

(11)
$$(X - \beta/\alpha)(X - \cos^2 \theta)^2 = X \sin^2 \theta \left\{ [1 + (r - 1)/2\alpha] X - \cos^2 \theta \right\}$$
, where $r = \bar{\rho}/\rho$, $\alpha = \frac{1}{2} [\gamma + 1 - (\gamma - 1)r]$, and $\beta = C_s^2/V_A^2 = 4\pi\gamma p/B^2$.

The density ratio is bounded by

(12)
$$1 < r < (\gamma + 1)/(\gamma - 1)$$
.

If the shock is normal to B (i.e., if $\theta = \pi/2$), then

(13)
$$U^2 = (r/\alpha) \left\{ C_{\bullet}^2 + V_A^2 \left[1 + (1 - \gamma/2)(r - 1) \right] \right\};$$

(14)
$$U/\bar{U} = \bar{B}/B = r;$$

(15)
$$\bar{V} = V$$
;

(16)
$$\bar{p} = p + (1 - r^{-1})\rho U^2 + (1 - r^2)B^2/2\mu$$
.

If $\theta = 0$, there are two possibilities: switch-on shocks, which require $\beta < 1$ and for which

(17)
$$U^2 = \tau V_A^2$$
:

(18)
$$\bar{U} = V_A^2/U$$
;

(19)
$$\bar{B}_{\perp}^{2} = 2B_{\parallel}^{2}(r-1)(\alpha-\beta);$$

(20)
$$\bar{V} = \bar{U}\bar{B}_{\perp}/B_{||};$$

(21)
$$\bar{p} = p + \rho U^2 (1 - \alpha + \beta) (1 - r^{-1}),$$

and acoustic (hydrodynamic) shocks, for which

(22)
$$U^2 = (r/\alpha)C_*^2$$
;

(23)
$$\tilde{U} = U/r$$
;

(24)
$$\tilde{V} = \tilde{B}_{\perp} = 0$$
;

(25)
$$\tilde{p} = p + \rho U^2 (1 - r^{-1}).$$

For acoustic shocks the specific volume and pressure are related by

(26)
$$\bar{v}/v = [(\gamma + 1)p + (\gamma - 1)\bar{p}]/[(\gamma - 1)p + (\gamma + 1)\bar{p}].$$

In terms of the upstream Mach number $M = U/C_s$,

(27)
$$\bar{\rho}/\rho = \upsilon/\bar{\upsilon} = U/\bar{U} = (\gamma + 1)M^2/[(\gamma - 1)M^2 + 2];$$

(28)
$$\bar{p}/p = (2\gamma M^2 - \gamma + 1)/(\gamma + 1)$$
;

(29)
$$\bar{T}/T = [(\gamma - 1)M^2 + 2](2\gamma M^2 - \gamma + 1)/(\gamma + 1)^2 M^2;$$

(30)
$$\bar{M}^2 = [(\gamma - 1)M^2 + 2]/[2\gamma M^2 - \gamma + 1].$$

The entropy change across the shock is

(31)
$$\Delta s \equiv \tilde{s} - s = c_v \ln[(\bar{p}/p)(\rho/\bar{\rho})^{\gamma}],$$

where $c_v = R/(\gamma - 1)m$ is the specific heat at constant volume; here R is the gas constant. In the weak-shock limit $(M \to 1)$,

(32)
$$\Delta s \to c_v \frac{2\gamma(\gamma-1)}{3(\gamma+1)} (M^2-1)^3 \approx \frac{16\gamma R}{3(\gamma+1)m} (M-1)^3$$
.

The radius at time t of a strong spherical blast wave resulting from the explosive release of energy E in a medium with uniform density ρ is

(33)
$$R_S = C_0 (Et^2/\rho)^{1/5}$$
,

where C_0 is a constant depending on γ . For $\gamma = 7/5$, $C_0 = 1.033$.

FUNDAMENTAL PLASMA PARAMETERS

All quantities are in Gaussian cgs units except temperature (T, T_e, T_i) expressed in eV and ion mass (m_i) expressed in units of the proton mass, $\mu = m_i/m_p$; Z is charge state: k is Boltzmann's constant: K is wavelength: γ is the adiabatic index: $\ln \Lambda$ is the Coulomb logarithm.

 $f_{ce} = \omega_{ce}/2\pi = 2.80 \times 10^6 B \,\mathrm{Hz}$

Frequencies

electron gyrofrequency

plasma skin depth

Debye length

| efection Shippednemel | Jee - wee/2" - 2.30 x 10 D 112 |
|--|---|
| | $\omega_{ce} = eB/m_ec = 1.76 \times 10^7 B \text{ rad/sec}$ |
| ion gyrofrequency | $f_{ci} = \omega_{ci}/2\pi = 1.52 \times 10^3 Z \mu^{-1} B \text{ Hz}$ |
| | $\omega_{ci} = eB/m_i c = 9.58 \times 10^3 Z \mu^{-1} B \text{ rad/sec}$ |
| electron plasma frequency | $f_{pe} = \omega_{pe}/2\pi = 8.98 \times 10^3 n_e^{1/2} \mathrm{Hz}$ |
| | $\omega_{pe} = (4\pi n_e e^2/m_e)^{1/2}$ |
| | $= 5.64 \times 10^4 n_e^{1/2} \text{rad/sec}$ |
| ion plasma frequency | $f_{pi} = \omega_{pi}/2\pi$ |
| | $= 2.10 \times 10^2 Z \mu^{-1/2} n_i^{1/2} \mathrm{Hz}$ |
| | $\omega_{pi} = (4\pi n_i Z^2 e^2/m_i)^{1/2}$ |
| | $= 1.32 \times 10^3 Z \mu^{-1/2} n_i^{1/2} \text{rad/sec}$ |
| electron trapping rate | $\nu_{Te} = (eKE/m_e)^{1/2}$ |
| • | $= 7.26 \times 10^8 K^{1/2} E^{1/2} \sec^{-1}$ |
| ion trapping rate | $\nu_{Ti} = (eKE/m_i)^{1/2}$ |
| | = $1.69 \times 10^7 K^{1/2} E^{1/2} \mu^{-1/2} sec^{-1}$ |
| electron collision rate | $\nu_e = 2.91 \times 10^{-6} n_e \ln \Lambda T_e^{-3/2} \text{ sec}^{-1}$ |
| ion collision rate | $\nu_i = 4.78 \times 10^{-8} n_i Z^2 \ln \Lambda T_i^{-3/2} \sec^{-1}$ |
| Lengths | |
| electron deBroglie length | $\lambda = \hbar/(m_e k T_e)^{1/2} = 2.76 \times 10^{-8} T_e^{-1/2} \text{ cm}$ |
| classical distance of minimum approach | $e^2/kT = 1.44 \times 10^{-7} T^{-1} \text{ cm}$ |
| electron gyroradius | $r_e = v_{Te}/\omega_{ce} = 2.38T_e^{1/2}B^{-1}$ cm |
| ion gyroradius | $r_i = v_{Ti}/\omega_{ci}$ |
| | = $1.02 \times 10^2 \mu^{1/2} Z^{-1} T_i^{1/2} B^{-1}$ cm |

 $c/\omega_{pe} = 5.31 \times 10^5 n_e^{-1/2} \text{ cm}$

 $= 7.43 \times 10^2 T^{1/2} n^{-1/2} \text{ cm}$

 $\lambda_D = (kT/4\pi ne^2)^{1/2}$

Velocities

electron thermal velocity

ion thermal velocity

ion sound velocity

Alfvén velocity

Dimensionless

(electron/proton mass ratio)^{1/2}

number of particles in Debye sphere

Alfvén velocity/speed of light

electron plasma/gyrofrequency

ion plasma/gyrofrequency ratio

thermal/magnetic energy ratio

magnetic/ion rest energy ratio

Miscellaneous

Bohm diffusion coefficient

transverse Spitzer resistivity

$$v_{Te} = (kT_e/m_e)^{1/2}$$

= $4.19 \times 10^7 T_e^{1/2}$ cm/sec
 $v_{Ti} = (kT_i/m_i)^{1/2}$
= $9.79 \times 10^5 \mu^{-1/2} T_i^{1/2}$ cm/sec
 $C_s = (\gamma Z k T_e/m_i)^{1/2}$

=
$$9.79 \times 10^5 (\gamma Z T_e/\mu)^{1/2}$$
 cm/sec
 $v_A = B/(4\pi n_i m_i)^{1/2}$

 $= 2.18 \times 10^{11} \mu^{-1/2} n_i^{-1/2} B \text{ cm/sec}$

$$(m_e/m_p)^{1/2} = 2.33 \times 10^{-2} = 1/42.9$$

 $(4\pi/3)n\lambda_D^3 = 1.72 \times 10^9 T^{3/2} n^{-1/2}$

$$v_A/c = 7.28\mu^{-1/2}n_i^{-1/2}B$$

 $\omega_{pe}/\omega_{ce} = 3.21 \times 10^{-3}n_e^{1/2}B^{-1}$

$$\omega_{pi}/\omega_{ci} = 0.137 \mu^{1/2} n_i^{1/2} B^{-1}$$

$$\beta = 8\pi n k T / B^2 = 4.03 \times 10^{-11} n T B^{-2}$$

$$B^2 / 8\pi n_i m_i c^2 = 26.5 \mu^{-1} n_i^{-1} B^2$$

$$D_B = (ckT/16eB)$$
= $6.25 \times 10^6 TB^{-1} \text{ cm}^2/\text{sec}$
 $\eta_{\perp} = 1.15 \times 10^{-14} Z \ln \Lambda T^{-3/2} \text{ sec}$
= $1.03 \times 10^{-2} Z \ln \Lambda T^{-3/2} \Omega \text{ cm}$

The anomalous collision rate due to low-frequency ion-sound turbulence is

$$\nu^* \approx \omega_{pe} \widetilde{W}/kT = 5.64 \times 10^4 n_e^{1/2} \widetilde{W}/kT \, {\rm sec}^{-1}$$
,

where \widetilde{W} is the total energy of waves with $\omega/K < v_{Ti}$. Magnetic pressure is given by

$$P_{\text{mag}} = B^2/8\pi = 3.98 \times 10^6 B^2 \text{ dynes/cm}^2 = 3.93 (B/B_0)^2 \text{ atm},$$

where $B_0 = 10 \, \text{kG} = 1 \, \text{T}$.

Detonation energy of 1 kiloton of high explosive is

$$W_{kT} = 10^{12} \text{ cal} = 4.2 \times \cdot 10^{19} \text{ erg.}$$

PLASMA DISPERSION FUNCTION

Definition¹⁶ (first form valid only for Im $\zeta > 0$):

$$Z(\zeta) = \pi^{-1/2} \int_{-\infty}^{+\infty} \frac{dt \exp\left(-t^2\right)}{t - \zeta} = 2i \exp\left(-\zeta^2\right) \int_{-\infty}^{i\zeta} dt \exp\left(-t^2\right).$$

Physically, $\zeta = x + iy$ is the ratio of wave phase velocity to thermal velocity.

Differential equation:

$$\frac{dZ}{d\zeta} = -2(1+\zeta Z), \ Z(0) = i\pi^{1/2}; \quad \frac{d^2Z}{d\zeta^2} + 2\zeta \frac{dZ}{d\zeta} + 2Z = 0.$$

Real argument (y = 0):

$$Z(x) = \exp\left(-x^2\right) \left(i\pi^{1/2} - 2\int_0^x dt \, \exp\left(t^2\right)\right).$$

Imaginary argument (x = 0):

$$Z(iy) = i\pi^{1/2} \exp(y^2) \left[1 - \operatorname{erf}(y)\right].$$

Power series (small argument):

$$Z(\zeta) = i\pi^{1/2} \exp\left(-\zeta^2\right) - 2\zeta \left(1 - 2\zeta^2/3 + 4\zeta^4/15 - 8\zeta^6/105 + \cdots\right).$$

Asymptotic series, $|\zeta| \gg 1$ (Ref. 17):

$$Z(\zeta) = i\pi^{1/2}\sigma \exp\left(-\zeta^2\right) - \zeta^{-1}\left(1 + 1/2\zeta^2 + 3/4\zeta^4 + 15/8\zeta^6 + \cdots\right),\,$$

where

$$\sigma = \begin{cases} 0 & y > |x|^{-1} \\ 1 & |y| < |x|^{-1} \\ 2 & y < -|x|^{-1} \end{cases}$$

Symmetry properties (the asterisk denotes complex conjugation):

$$Z(\zeta^*) = -[Z(-\zeta)]^*;$$

$$Z(\zeta^*) = [Z(\zeta)]^* + 2i\pi^{1/2} \exp[-(\zeta^*)^2] \quad (y > 0).$$

Two-pole approximations¹⁸ (good for ζ in upper half plane except when $y < \pi^{1/2} x^2 \exp(-x^2)$, $x \gg 1$):

$$Z(\zeta) \approx \frac{0.50 + 0.81i}{a - \zeta} - \frac{0.50 - 0.81i}{a^* + \zeta}, \quad a = 0.51 - 0.81i;$$

$$Z'(\zeta) \approx \frac{0.50 + 0.96i}{(b - \zeta)^2} + \frac{0.50 - 0.96i}{(b^* + \zeta)^2}, \ b = 0.48 - 0.91i.$$

COLLISIONS AND TRANSPORT

Temperatures are in eV; the corresponding value of Boltzmann's constant is $k = 1.60 \times 10^{-12} \, \mathrm{erg/eV}$; masses μ , μ' are in units of the proton mass; $e_{\alpha} = Z_{\alpha}e$ is the charge of species α . All other units are cgs except where noted.

Relaxation Rates

Rates are associated with four relaxation processes arising from the interaction of test particles (labeled α) streaming with velocity \mathbf{v}_{α} through a background of field particles (labeled β):

slowing down
$$\frac{d\mathbf{v}_{\alpha}}{dt} = -\nu_{*}^{\alpha/\beta}\mathbf{v}_{\alpha}$$
transverse diffusion
$$\frac{d}{dt}(\mathbf{v}_{\alpha} - \bar{\mathbf{v}}_{\alpha})_{\perp}^{2} = \nu_{\perp}^{\alpha/\beta}v_{\alpha}^{2}$$
parallel diffusion
$$\frac{d}{dt}(\mathbf{v}_{\alpha} - \bar{\mathbf{v}}_{\alpha})_{\parallel}^{2} = \nu_{\parallel}^{\alpha/\beta}v_{\alpha}^{2}$$
energy loss
$$\frac{d}{dt}v_{\alpha}^{2} = -\nu_{\epsilon}^{\alpha/\beta}v_{\alpha}^{2},$$

where the averages are performed over an ensemble of test particles and a Maxwellian field particle distribution. The exact formulas may be written¹⁹

$$\begin{split} \nu_s^{\alpha/\beta} &= (1 + m_\alpha/m_\beta) \psi(x^{\alpha/\beta}) \nu_0^{\alpha/\beta}; \\ \nu_\perp^{\alpha/\beta} &= 2 \left[(1 - 1/2x^{\alpha/\beta}) \psi(x^{\alpha/\beta}) + \psi'(x^{\alpha/\beta}) \right] \nu_0^{\alpha/\beta}; \\ \nu_\parallel^{\alpha/\beta} &= \left[\psi(x^{\alpha/\beta})/x^{\alpha/\beta} \right] \nu_0^{\alpha/\beta}; \\ \nu_\epsilon^{\alpha/\beta} &= 2 \left[(m_\alpha/m_\beta) \psi(x^{\alpha/\beta}) - \psi'(x^{\alpha/\beta}) \right] \nu_0^{\alpha/\beta}, \end{split}$$

where

$$\nu_0^{\alpha/\beta} = 4\pi e_{\alpha}^{2} e_{\beta}^{2} \lambda_{\alpha\beta} n_{\beta} / m_{\alpha}^{2} v_{\alpha}^{3}; \qquad x^{\alpha/\beta} = m_{\beta} v_{\alpha}^{2} / 2kT_{\beta};$$

$$\psi(x) = \frac{2}{\sqrt{\pi}} \int_{0}^{x} dt \, t^{1/2} e^{-t}; \quad \psi'(x) = \frac{d\psi}{dx},$$

and $\lambda_{\alpha\beta} = \ln \Lambda_{\alpha\beta}$ is the Coulomb logarithm (see below). Limiting forms of ν_a , ν_{\perp} and ν_{\parallel} are given in the following table. All the expressions shown

have units cm³ sec⁻¹. Test particle energy ϵ and field particle temperature T are both in eV; $\mu = m_i/m_p$ where m_p is the proton mass: Z is ion charge state; in electron-electron and ion-ion encounters, field particle quantities are distinguished by a prime. The two expressions given below for each rate hold for very slow $(x^{\alpha/3} \ll 1)$ and very fast $(x^{\alpha/3} \gg 1)$ test particles, respectively.

Electron-electron
$$v_{s}^{e/e'}/n_{e'}\lambda_{ee'}\approx 5.8\times 10^{-6}T^{-3/2} \longrightarrow 7.7\times 10^{-6}e^{-3/2}$$

$$v_{\perp}^{e/e'}/n_{e'}\lambda_{ee'}\approx 5.8\times 10^{-6}T^{-1/2}e^{-1} \longrightarrow 7.7\times 10^{-6}e^{-3/2}$$

$$v_{\parallel}^{e/e'}/n_{e'}\lambda_{ee'}\approx 2.9\times 10^{-6}T^{-1/2}e^{-1} \longrightarrow 3.9\times 10^{-6}Te^{-5/2}$$
Electron-ion
$$v_{\parallel}^{e/i}/n_{i}Z^{2}\lambda_{ei}\approx 0.23\mu^{3/2}T^{-3/2} \longrightarrow 3.9\times 10^{-6}e^{-3/2}$$

$$v_{\perp}^{e/i}/n_{i}Z^{2}\lambda_{ei}\approx 2.5\times 10^{-4}\mu^{1/2}T^{-1/2}e^{-1} \longrightarrow 7.7\times 10^{-6}e^{-3/2}$$

$$v_{\parallel}^{e/i}/n_{i}Z^{2}\lambda_{ei}\approx 1.2\times 10^{-4}\mu^{1/2}T^{-1/2}e^{-1} \longrightarrow 2.1\times 10^{-9}\mu^{-1}Te^{-5/2}$$
Ion-electron
$$v_{\parallel}^{e/e}/n_{e}Z^{2}\lambda_{ie}\approx 1.6\times 10^{-9}\mu^{-1}T^{-3/2} \longrightarrow 1.7\times 10^{-4}\mu^{1/2}e^{-3/2}$$

$$v_{\parallel}^{i/e}/n_{e}Z^{2}\lambda_{ie}\approx 3.2\times 10^{-9}\mu^{-1}T^{-1/2}e^{-1} \longrightarrow 1.8\times 10^{-7}\mu^{-1/2}e^{-3/2}$$

$$v_{\parallel}^{i/e}/n_{e}Z^{2}\lambda_{ie}\approx 1.6\times 10^{-9}\mu^{-1}T^{-1/2}e^{-1} \longrightarrow 1.7\times 10^{-4}\mu^{1/2}Te^{-5/2}$$
Ion-ion
$$\frac{v_{\parallel}^{i/i'}}{n_{i'}Z^{2}Z^{2}\lambda_{ii'}}\approx 6.8\times 10^{-8}\frac{\mu'^{1/2}}{\mu}\left(1+\frac{\mu'}{\mu}\right)T^{-3/2} \longrightarrow 9.0\times 10^{-8}\left(\frac{1}{\mu}+\frac{1}{\mu'}\right)\frac{\mu^{1/2}}{e^{3/2}}$$

$$\frac{v_{\parallel}^{i/i'}}{n_{i'}Z^{2}Z^{2/2}\lambda_{ii'}}\approx 1.4\times 10^{-7}\mu'^{1/2}\mu^{-1}T^{-1/2}e^{-1} \longrightarrow 1.8\times 10^{-7}\mu^{-1/2}e^{-3/2}$$

$$\longrightarrow 9.0\times 10^{-8}\left(\frac{1}{\mu}+\frac{1}{\mu'}\right)\frac{\mu^{1/2}}{e^{3/2}}$$

$$\longrightarrow 9.0\times 10^{-8}\mu^{1/2}\mu'^{-1}Te^{-5/2}$$

$$\longrightarrow 9.0\times 10^{-8}\mu^{1/2}\mu'^{-1}Te^{-5/2}$$

In the same limits, the energy transfer rate follows from the identity

$$\nu_{\epsilon} = 2\nu_s - \nu_{\perp} - \nu_{\parallel},$$

except for the case of fast electrons or fast ions scattered by ions, where the leading terms cancel. Then the appropriate forms are

$$\nu_{\epsilon}^{e/i} \longrightarrow 4.2 \times 10^{-9} n_i Z^2 \lambda_{ei}$$

$$\left[\epsilon^{-3/2} \mu^{-1} - 8.9 \times 10^4 (\mu/T)^{1/2} \epsilon^{-1} \exp(-1836 \mu \epsilon/T) \right] \sec^{-1}$$

and

$$\nu_{\epsilon}^{i/i'} \longrightarrow 1.8 \times 10^{-7} n_{i'} Z^2 Z'^2 \lambda_{ii'}$$

$$\left[\epsilon^{-3/2} \mu^{1/2} / \mu' - 1.1 (\mu'/T)^{1/2} \epsilon^{-1} \exp(-\mu' \epsilon/T) \right] \sec^{-1}.$$

In general, the energy transfer rate $\nu_e^{\alpha/\beta}$ is positive for $\epsilon > \epsilon_\alpha^*$ and negative for $\epsilon < \epsilon_{\alpha}^*$, where $x^* = (m_\beta/m_\alpha)\epsilon_\alpha^*/T_\beta$ is the solution of $\psi'(x^*) = (m_\alpha/m_\beta)\psi(x^*)$. The ratio $\epsilon_\alpha^*/T_\beta$ is given for a number of specific α , β in the following table:

| α/β | i/e | e/e, i/i | e/p | e/D | e/T, e/He3 | e/He4 |
|---|-----|----------|----------------------|----------------------|----------------------|----------------------|
| $\frac{\epsilon_{\alpha}^{*}}{T_{\beta}}$ | 1.5 | 0.98 | 4.8×10^{-3} | 2.6×10^{-3} | 1.8×10^{-3} | 1.4×10^{-3} |

When both species are near Maxwellian, with $T_i \lesssim T_e$, there are just two characteristic collision rates. For Z = 1,

$$\nu_e = 2.9 \times 10^{-6} n \lambda T_e^{-3/2} \text{ sec}^{-1};$$

$$\nu_i = 4.8 \times 10^{-6} n \lambda T_i^{-3/2} \mu^{-1/2} \text{ sec}^{-1}.$$

Temperature Isotropization

Isotropization is described by

$$\frac{dT_{\perp}}{dt} = -\frac{1}{2}\frac{dT_{\parallel}}{dt} = -\nu_T^{\alpha}(T_{\perp} - T_{\parallel}),$$

where, if $A \equiv T_{\perp}/T_{\parallel} - 1 > 0$,

$$\nu_T^{\alpha} = \frac{2\sqrt{\pi}e_{\alpha}^2 e_{\beta}^2 n_{\alpha} \lambda_{\alpha\beta}}{m_{\alpha}^{1/2} (kT_{||})^{3/2}} A^{-2} \left[-3 + (A+3) \frac{\tan^{-1}(A^{1/2})}{A^{1/2}} \right].$$

If A < 0, $\tan^{-1}(A^{1/2})/A^{1/2}$ is replaced by $\tanh^{-1}(-A)^{1/2}/(-A)^{1/2}$. For $T_{\perp} \approx T_{\parallel} \equiv T$,

$$\nu_T^e = 8.2 \times 10^{-7} n \lambda T^{-3/2} \text{ sec}^{-1};$$

$$\nu_T^i = 1.9 \times 10^{-8} n \lambda Z^2 \mu^{-1/2} T^{-3/2} \text{ sec}^{-1}.$$

Thermal Equilibration

If the components of a plasma have different temperatures, but no relative drift, equilibration is described by

$$\frac{dT_{\alpha}}{dt} = \sum_{\beta} \bar{\nu}_{\epsilon}^{\alpha/\beta} (T_{\beta} - T_{\alpha}).$$

where

$$\bar{\nu}_{\epsilon}^{\alpha/3} = 1.8 \times 10^{-19} \frac{(m_{\alpha} m_{\beta})^{1/2} Z_{\alpha}^{2} Z_{\beta}^{2} n_{\beta} \lambda_{\alpha\beta}}{(m_{\alpha} T_{\beta} + m_{\beta} T_{\alpha})^{3/2}} \sec^{-1}.$$

For electrons and ions with $T_e \approx T_i \equiv T$, this implies

$$\bar{\nu}_{\epsilon}^{e/i}/n_i = \bar{\nu}_{\epsilon}^{i/e}/n_e = 3.2 \times 10^{-9} Z^2 \lambda/\mu T^{3/2} \text{cm}^3 \text{sec}^{-1}.$$

Coulomb Logarithm

For test particles of mass m_{α} and charge $e_{\alpha} = Z_{\alpha}e$ scattering off field particles of mass m_{β} and charge $e_{\beta} = Z_{\beta}e$, the Coulomb logarithm is defined as $\lambda = \ln \Lambda \equiv \ln(r_{\max}/r_{\min})$. Here r_{\min} is the larger of $e_{\alpha}e_{\beta}/m_{\alpha\beta}\bar{u}^2$ and $\hbar/2m_{\alpha\beta}\bar{u}$, averaged over both particle velocity distributions, where $m_{\alpha\beta} = m_{\alpha}m_{\beta}/(m_{\alpha}+m_{\beta})$ and $u = v_{\alpha}-v_{\beta}$; $r_{\max} = (4\pi\sum_{\alpha}n_{\gamma}e_{\gamma}^{2}/kT_{\gamma})^{-1/2}$, where the summation extends over all species γ for which $\bar{u}^2 < v_{T\gamma}^2 = kT_{\gamma}/m_{\gamma}$. If this inequality cannot be satisfied, or if either $\bar{u}\omega_{c\alpha}^{-1} < r_{\max}$ or $\bar{u}\omega_{c\beta}^{-1} < r_{\max}$, the theory breaks down. Typically $\lambda \approx 10$ –20. Corrections to the transport coefficients are $O(\lambda^{-1})$; hence the theory is good only to $\sim 10\%$ and fails when $\lambda \sim 1$.

The following cases are of particular interest:

(a) Thermal electron-electron collisions

$$\lambda_{ee} = 23 - \ln(n_e^{1/2} T_e^{-3/2}), \qquad T_e \lesssim 10 \,\text{eV};$$

= $24 - \ln(n_e^{1/2} T_e^{-1}), \qquad T_e \gtrsim 10 \,\text{eV}.$

(b) Electron-ion collisions

$$\begin{split} \lambda_{ei} &= \lambda_{ie} = 23 - \ln\left(n_e^{1/2} Z T_e^{-3/2}\right), & T_i m_e/m_i < T_e < 10 Z^2 \, \text{eV}; \\ &= 24 - \ln\left(n_e^{1/2} T_e^{-1}\right), & T_i m_e/m_i < 10 Z^2 \, \text{eV} < T_e \\ &= 30 - \ln\left(n_i^{1/2} T_i^{-3/2} Z^2 \mu^{-1}\right), & T_e < T_i Z m_e/m_i. \end{split}$$

(c) Mixed ion-ion collisions

$$\lambda_{ii'} = \lambda_{i'i} = 23 - \ln \left[\frac{ZZ'(\mu + \mu')}{\mu T_{i'} + \mu' T_{i}} \left(\frac{n_{i}Z^{2}}{T_{i}} + \frac{n_{i'}Z'^{2}}{T_{i'}} \right)^{1/2} \right].$$

(d) Counterstreaming ions (relative velocity $v_D = \beta_D c$) in the presence of warm electrons, kT_i/m_i , $kT_{i'}/m_{i'} < v_D^2 < kT_e/m_e$

$$\lambda_{ii'} = \lambda_{i'i} = 35 - \ln \left[\frac{ZZ'(\mu + \mu')}{\mu \mu' \beta_D^2} \left(\frac{n_e}{T_e} \right)^{1/2} \right].$$

Fokker-Planck Equation

$$\frac{Df^{\alpha}}{Dt} \equiv \frac{\partial f^{\alpha}}{\partial t} + \mathbf{v} \cdot \nabla f^{\alpha} + \mathbf{F} \cdot \nabla_{\mathbf{v}} f^{\alpha} = \left(\frac{\partial f^{\alpha}}{\partial t}\right)_{\text{coll}},$$

where F is an external force field. The general form of the collision integral is $(\partial f^{\alpha}/\partial t)_{\text{coll}} = -\sum_{\beta} \nabla_{\mathbf{v}} \cdot \mathbf{J}^{\alpha/\beta}$, with

$$\mathbf{J}^{\alpha/\beta} = 2\pi \lambda_{\alpha\beta} \frac{e_{\alpha}^{2} e_{\beta}^{2}}{m_{\alpha}} \int d^{3}v'(u^{2}I - uu)u^{-3}$$
$$\cdot \left\{ \frac{1}{m_{\beta}} f^{\alpha}(\mathbf{v}) \nabla_{\mathbf{v}'} f^{\beta}(\mathbf{v}') - \frac{1}{m_{\alpha}} f^{\beta}(\mathbf{v}') \nabla_{\mathbf{v}} f^{\alpha}(\mathbf{v}) \right\}$$

(Landau form) where u = v' - v and l is the unit dyad, or alternatively,

$$\mathbf{J}^{\alpha/\beta} = 4\pi\lambda_{\alpha\beta}\frac{{e_{\alpha}}^2{e_{\beta}}^2}{{m_{\alpha}}^2}\left\{f^{\alpha}(\mathbf{v})\nabla_{\mathbf{v}}H(\mathbf{v}) - \frac{1}{2}\nabla_{\mathbf{v}}\cdot\left[f^{\alpha}(\mathbf{v})\nabla_{\mathbf{v}}\nabla_{\mathbf{v}}G(\mathbf{v})\right]\right\},$$

where the Rosenbluth potentials are

$$G(\mathbf{v}) = \int f^{3}(\mathbf{v}')ud^{3}v'$$

$$H(\mathbf{v}) = \left(1 + \frac{m_{\alpha}}{m_{\beta}}\right) \int f^{\beta}(\mathbf{v}') u^{-1} d^{3}v'.$$

If species α is a weak beam (number and energy density small compared with background) streaming through a Maxwellian plasma, then

$$\mathbf{J}^{\alpha/3} = -\nu_s^{\alpha/3} \mathbf{v} f^{\alpha} - \frac{1}{2} \nu_{\perp}^{\alpha/3} v^2 \nabla_{\mathbf{v}} f^{\alpha} + \frac{1}{2} (\nu_{\perp}^{\alpha/3} - \nu_{\parallel}^{\alpha/3}) \mathbf{v} \mathbf{v} \cdot \nabla_{\mathbf{v}} f^{\alpha}.$$

B-G-K Collision Operator

For distribution functions with no large gradients in velocity space, the Fokker-Planck collision terms can be approximated according to

$$\frac{Df_e}{Dt} = \nu_{ee}(F_e - f_e) + \nu_{ei}(\bar{F}_e - f_e);$$

$$\frac{Df_i}{Dt} = \nu_{ie}(\bar{F}_i - f_i) + \nu_{ii}(F_i - f_i).$$

The respective slowing-down rates $\nu_{\bullet}^{\alpha/\beta}$ given in the Relaxation Rate section above can be used for $\nu_{\alpha\beta}$, assuming slow ions and fast electrons, with ϵ replaced by T_{α} . (For ν_{ee} and ν_{ii} , one can equally well use ν_{\perp} , and the result is insensitive to whether the slow- or fast-test-particle limit is employed.) The Maxwellians F_{α} and \bar{F}_{α} are given by

$$F_{\alpha} = n_{\alpha} \left(\frac{m_{\alpha}}{2\pi k T_{\alpha}} \right)^{3/2} \exp \left\{ -\left[\frac{m_{\alpha} (\mathbf{v} - \mathbf{v}_{\alpha})^{2}}{2k T_{\alpha}} \right] \right\};$$

$$\bar{F}_{\alpha} = n_{\alpha} \left(\frac{m_{\alpha}}{2\pi k \bar{T}_{\alpha}} \right)^{3/2} \exp \left\{ - \left[\frac{m_{\alpha} (\mathbf{v} - \bar{\mathbf{v}}_{\alpha})^{2}}{2k \bar{T}_{\alpha}} \right] \right\},$$

where n_{α} , \mathbf{v}_{α} and T_{α} are the number density, mean drift velocity, and effective temperature obtained by taking moments of f_{α} . Some latitude in the definition of \bar{T}_{α} and $\bar{\mathbf{v}}_{\alpha}$ is possible;²⁰ one choice is $\bar{T}_{e} = T_{i}$, $\bar{T}_{i} = T_{e}$, $\bar{\mathbf{v}}_{e} = \mathbf{v}_{i}$, $\bar{\mathbf{v}}_{i} = \mathbf{v}_{e}$.

Transport Coefficients

Transport equations for a multispecies plasma:

$$\frac{d^{\alpha} n_{\alpha}}{dt} + n_{\alpha} \nabla \cdot \mathbf{v}_{\alpha} = 0;$$

$$m_{\alpha}n_{\alpha}\frac{d^{\alpha}v_{\alpha}}{dt}=-\nabla p_{\alpha}-\nabla\cdot P_{\alpha}+Z_{\alpha}en_{\alpha}\left[\mathbf{E}+\frac{1}{c}\mathbf{v}_{\alpha}\times\mathbf{B}\right]+\mathbf{R}_{\alpha};$$

$$\frac{3}{2}n_{\alpha}\frac{d^{\alpha}kT_{\alpha}}{dt} + p_{\alpha}\nabla \cdot \mathbf{v}_{\alpha} = -\nabla \cdot \mathbf{q}_{\alpha} - P_{\alpha} : \nabla \mathbf{v}_{\alpha} + Q_{\alpha}.$$

Here $d^{\alpha}/dt \equiv \partial/\partial t + \mathbf{v}_{\alpha} \cdot \nabla$: $p_{\alpha} = n_{\alpha}kT_{\alpha}$, where k is Boltzmann's constant; $\mathbf{R}_{\alpha} = \sum_{\beta} \mathbf{R}_{\alpha\beta}$ and $Q_{\alpha} = \sum_{\beta} Q_{\alpha\beta}$, where $\mathbf{R}_{\alpha\beta}$ and $Q_{\alpha\beta}$ are respectively the momentum and energy gained by the α th species through collisions with the β th: P_{α} is the stress tensor: and \mathbf{q}_{α} is the heat flow.

The transport coefficients in a simple two-component plasma (electrons and singly charged ions) are tabulated below. Here || and \perp refer to the direction of the magnetic field B = bB; $u = v_e - v_i$ is the relative streaming velocity; $n_e = n_i \equiv n$; j = -neu is the current; $\omega_{ce} = 1.76 \times 10^7 B \text{ sec}^{-1}$ and $\omega_{ci} = (m_e/m_i)\omega_{ce}$ are the electron and ion gyrofrequencies, respectively; and the basic collisional times are taken to be

$$\tau_e = \frac{3\sqrt{m_e(kT_e)^{3/2}}}{4\sqrt{2\pi} n\lambda e^4} = 3.44 \times 10^5 \frac{T_e^{3/2}}{n\lambda} \sec,$$

where λ is the Coulomb logarithm, and

$$\tau_i = \frac{3\sqrt{m_i}(kT_i)^{3/2}}{4\sqrt{\pi}n\,\lambda e^4} = 2.09 \times 10^7 \frac{T_i^{3/2}}{n\lambda} \mu^{1/2} \text{ sec.}$$

In the limit of large fields ($\omega_{c\alpha}\tau_{\alpha}\gg 1$, $\alpha=i,e$) the transport processes may be summarized as follows:²¹

momentum transfer $R_{ij} = -R_{ij} \equiv R = R_{ii} + R_{T}$: $\mathbf{R}_{\mathbf{u}} = ne(\mathbf{j}_{\parallel}/\sigma_{\parallel} + \mathbf{j}_{\perp}/\sigma_{\perp});$ frictional force $\sigma_{\parallel} = 2.0\sigma_{\perp} = 2.0 \frac{ne^2 \tau_e}{m};$ electrical conductivities $\mathbf{R}_T = -0.71 n \nabla_{\parallel}(kT_e) - \frac{3n}{2\omega_{\perp}\tau} \mathbf{b} \times \nabla_{\perp}(kT_e);$ thermal force $Q_i = \frac{3m_e}{m_i} \frac{nk}{\pi} (T_e - T_i);$ ion heating $Q_e = -Q_i - \mathbf{R} \cdot \mathbf{u};$ electron heating $\mathbf{q}_{i} = -\kappa_{\parallel}^{i} \nabla_{\parallel}(kT_{i}) - \kappa_{\perp}^{i} \nabla_{\perp}(kT_{i}) + \kappa_{\wedge}^{i} \mathbf{b} \times \nabla_{\perp}(kT_{i});$ ion heat flux $\kappa_{\parallel}^{i} = 3.9 \frac{nkT_{i}\tau_{i}}{m_{i}}; \quad \kappa_{\perp}^{i} = \frac{2nkT_{i}}{m_{i}\omega^{2}\tau_{i}}; \quad \kappa_{\wedge}^{i} = \frac{5nkT_{i}}{2m_{i}\omega_{ci}};$ ion thermal conductivities $q_e = q_u^e + q_T^e$; electron heat flux $\mathbf{q}_{\mathbf{u}}^{e} = 0.71nkT_{e}\mathbf{u}_{\parallel} + \frac{3nkT_{e}}{2m_{e}\tau_{e}}\mathbf{b} \times \mathbf{u}_{\perp};$ frictional heat flux

thermal gradient heat flux
$$\begin{aligned} \mathbf{q}_{T}^{e} &= -\kappa_{\parallel}^{e} \nabla_{\parallel}(kT_{e}) - \kappa_{\perp}^{e} \nabla_{\perp}(kT_{e}) - \kappa_{\wedge}^{e} \mathbf{b} \times \nabla_{\perp}(kT_{e}); \\ \text{heat flux} \end{aligned}$$
 electron thermal conductivities
$$\kappa_{\parallel}^{e} &= 3.2 \frac{nkT_{e}\tau_{e}}{m_{e}}; \quad \kappa_{\perp}^{e} = 4.7 \frac{nkT_{e}}{m_{e}\omega_{ce}^{2}\tau_{e}}; \quad \kappa_{\wedge}^{e} = \frac{5nkT_{e}}{2m_{e}\omega_{ce}}; \end{aligned}$$
 stress tensor (both species)
$$P_{xx} = -\frac{\eta_{0}}{2}(W_{xx} + W_{yy}) - \frac{\eta_{1}}{2}(W_{xx} - W_{yy}) - \eta_{3}W_{xy}; \end{aligned}$$

$$P_{yy} = -\frac{\eta_{0}}{2}(W_{xx} + W_{yy}) + \frac{\eta_{1}}{2}(W_{xx} - W_{yy}) + \eta_{3}W_{xy};$$

$$P_{xy} = P_{yx} = -\eta_{1}W_{xy} + \frac{\eta_{3}}{2}(W_{xx} - W_{yy});$$

$$P_{xz} = P_{zx} = -\eta_{2}W_{xz} - \eta_{4}W_{yz};$$

$$P_{yz} = P_{zy} = -\eta_{2}W_{yz} + \eta_{4}W_{xz};$$

$$P_{zz} = -\eta_{0}W_{zz}$$

(here the z axis is defined parallel to B);

For both species the rate-of-strain tensor is defined as

$$W_{jk} = \frac{\partial v_j}{\partial x_k} + \frac{\partial v_k}{\partial x_j} - \frac{2}{3} \delta_{jk} \nabla \cdot \mathbf{v}.$$

When B = 0 the following simplifications occur:

$$\mathbf{R}_{\mathbf{u}} = ne\mathbf{j}/\sigma_{||}; \quad \mathbf{R}_{T} = -0.71n\nabla(kT_{e}); \quad \mathbf{q}_{i} = -\kappa_{||}^{i}\nabla(kT_{i});$$

$$\mathbf{q}_{\mathbf{u}}^{e} = 0.71nkT_{e}\mathbf{u}; \quad \mathbf{q}_{T}^{e} = -\kappa_{||}^{e}\nabla(kT_{e}); \quad P_{jk} = -\eta_{0}W_{jk}.$$

For $\omega_{ce}\tau_e\gg 1\gg \omega_{ci}\tau_i$, the electrons obey the high-field expressions and the ions obey the zero-field expressions.

Collisional transport theory is applicable when (1) macroscopic time rates of change satisfy $d/dt \ll 1/\tau$, where τ is the longest collisional time scale, and (in the absence of a magnetic field) (2) macroscopic length scales L satisfy $L \gg l$, where $l = \bar{v}\tau$ is the mean free path. In a strong field, $\omega_{ce}\tau \gg 1$, condition (2) is replaced by $L_{\parallel} \gg l$ and $L_{\perp} \gg \sqrt{lr_e}$ ($L_{\perp} \gg r_e$ in a uniform field),

where L_{\parallel} is a macroscopic scale parallel to the field B and L_{\perp} is the smaller of $B/|\nabla_{\perp}B|$ and the transverse plasma dimension. In addition, the standard transport coefficients are valid only when (3) the Coulomb logarithm satisfies $\lambda \gg 1$: (4) the electron gyroradius satisfies $r_e \gg \lambda_D$, or $8\pi n_e m_e c^2 \gg B^2$; (5) relative drifts $u = v_{\alpha} - v_{\beta}$ between two species are small compared with the thermal velocities, i.e., $u^2 \ll kT_{\alpha}/m_{\alpha}$, kT_{β}/m_{β} ; and (6) anomalous transport processes owing to microinstabilities are negligible.

Weakly Ionized Plasmas

Collision frequency for scattering of charged particles of species α by neutrals is

$$\nu_{\alpha} = n_0 \sigma_s^{\alpha/0} (kT_{\alpha}/m_{\alpha})^{1/2},$$

where n_0 is the neutral density and $\sigma_s^{\alpha/0}$ is the cross section, typically $\sim 5 \times 10^{-15}$ cm² and weakly dependent on temperature.

When the system is small compared with a Debye length, $L \ll \lambda_D$, the charged particle diffusion coefficients are

$$D_{\alpha} = kT_{\alpha}/m_{\alpha}\nu_{\alpha},$$

In the opposite limit, both species diffuse at the ambipolar rate

$$D_{A} = \frac{\mu_{i} D_{e} - \mu_{e} D_{i}}{\mu_{i} - \mu_{e}} = \frac{(T_{i} + T_{e}) D_{i} D_{e}}{T_{i} D_{e} + T_{e} D_{i}},$$

where $\mu_{\alpha}=e_{\alpha}/m_{\alpha}\nu_{\alpha}$ is the mobility. The conductivity σ_{α} satisfies $\sigma_{\alpha}=n_{\alpha}e_{\alpha}\mu_{\alpha}$.

In the presence of a magnetic field B the scalars μ and σ become tensors,

$$\mathbf{J}^{\alpha} = \boldsymbol{\sigma}^{\alpha} \cdot \mathbf{E} = \sigma_{\parallel}^{\alpha} \mathbf{E}_{\parallel} + \sigma_{\perp}^{\alpha} \mathbf{E}_{\perp} + \sigma_{\wedge}^{\alpha} \mathbf{E} \times \mathbf{b},$$

where b = B/B and

$$\sigma_{\parallel}^{\alpha} = n_{\alpha} e_{\alpha}^{2} / m_{\alpha} \nu_{\alpha};$$

$$\sigma_{\perp}^{\alpha} = \sigma_{\parallel}^{\alpha} \nu_{\alpha}^{2} / (\nu_{\alpha}^{2} + \omega_{c\alpha}^{2});$$

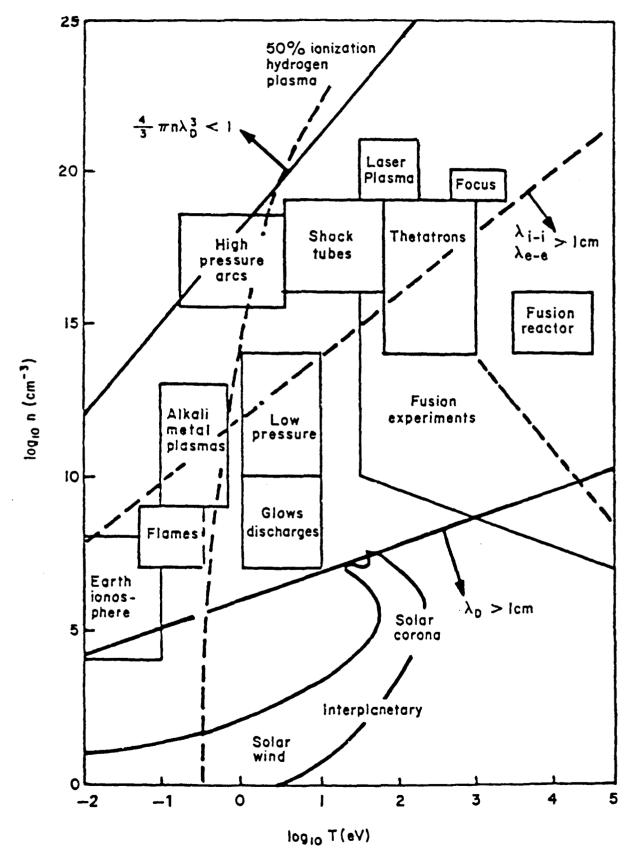
$$\sigma_{\wedge}^{\alpha} = \sigma_{\parallel}^{\alpha} \nu_{\alpha} \omega_{c\alpha} / (\nu_{\alpha}^{2} + \omega_{c\alpha}^{2}).$$

Here σ_{\perp} and σ_{\wedge} are the Pedersen and Hall conductivities, respectively.

APPROXIMATE MAGNITUDES IN SOME TYPICAL PLASMAS

| Plasma Type | $n \text{ cm}^{-3}$ | T eV | $\omega_{pe} \sec^{-1}$ | $\lambda_{\mathcal{D}}$ cm | $n\lambda_D^3$ | $v_{ei} \mathrm{sec}^{-1}$ |
|---------------------------------|---------------------|-----------------|--------------------------|----------------------------|---------------------|-----------------------------|
| Interstellar gas | 1 | 1 | 6 × 10 ⁴ | 7×10^2 | 4 × 10 ⁸ | 7×10^{-5} |
| Gaseous nebula | 10 ³ | 1 | 2×10^6 | 20 | 10 ⁷ | 6×10^{-2} |
| Solar Corona | 10 ⁶ | 10 ² | 6×10^{7} | 7 | 4×10^3 | 6×10^{-2} |
| Diffuse hot plasma | 1012 | 10 ² | 6×10^{10} | 7×10^{-3} | 4×10^5 | 40 |
| Solar atmosphere, gas discharge | 1014 | 1 | 6 × 10 ¹¹ | 7×10^{-5} | 40 | 2×10^9 |
| Warm plasma | 1014 | 10 | 6×10^{11} | 2×10^{-4} | 10 ³ | 107 |
| Hot plasma | 1014 | 10 ² | 6×10^{11} | 7×10^{-4} | 4×10^4 | 4×10^6 |
| Thermonuclear plasma | 10 ¹⁵ | 104 | 2×10^{12} | 2×10^{-3} | 107 | 5 × 10 ⁴ |
| Theta pinch | 10 ¹⁶ | 10 ² | 6×10^{12} | 7×10^{-5} | 4×10^3 | 3×10^{8} |
| Dense hot plasma | 10 ¹⁸ | 10 ² | 6×10^{13} | 7×10^{-6} | 4×10^2 | 2×10^{10} |
| Laser Plasma | 10 ²⁰ | 10 ² | 6 × 10 ¹⁴ | 7×10^{-7} | 40 | 2×10^{12} |

The diagram (facing) gives comparable information in graphical form.²²



IONOSPHERIC PARAMETERS²³

The following tables give average nighttime values. Where two numbers are entered, the first refers to the lower and the second to the upper portion of the layer.

| Quantity | E Region | F Region |
|---|---|---|
| Altitude (km) | 90-160 | 160-500 |
| Number density (m ⁻³) | $1.5 \times 10^{10} - 3.0 \times 10^{10}$ | $5 \times 10^{10} - 2 \times 10^{11}$ |
| Height-integrated number density (m ⁻²) | 9 × 10 ¹⁴ | 4.5×10^{15} |
| Ion-neutral collision frequency (sec ⁻¹) | $2\times10^3-10^2$ | 0.5-0.05 |
| Ion gyro-/collision frequency ratio κ_i | 0.09-2.0 | $4.6 \times 10^2 - 5.0 \times 10^3$ |
| Ion Pederson factor $\kappa_i/(1+{\kappa_i}^2)$ | 0.09~0.5 | $2.2 \times 10^{-3} - 2 \times 10^{-4}$ |
| Ion Hall factor $\kappa_i^2/(1+\kappa_i^2)$ | 8 × 10 ⁻⁴ -0.8 | 1.0 |
| Electron-neutral collision frequency | $1.5 \times 10^4 - 9.0 \times 10^2$ | 80-10 |
| Electron gyro-/collision frequency ratio κ_e | $4.1 \times 10^2 - 6.9 \times 10^3$ | $7.8 \times 10^4 - 6.2 \times 10^5$ |
| Electron Pedersen factor $\kappa_e/(1+\kappa_e^2)$ | 2.7×10^{-3} - 1.5×10^{-4} | $10^{-5} - 1.5 \times 10^{-6}$ |
| Electron Hall factor $\kappa_e^2/(1+\kappa_e^2)$ | 1.0 | 1.0 |
| Mean molecular weight | 28-26 | 22-16 |
| Ion gyrofrequency (sec ⁻¹) | 180-190 | 230-300 |
| Neutral diffusion coefficient (m ² sec ⁻¹) | $30-5\times10^3$ | 10 ⁵ |

The terrestrial magnetic field in the lower ionosphere at equatorial lattitudes is approximately $B_0 = 0.35 \times 10^{-4}$ tesla. The earth's radius is $R_E = 6371$ km.

SOLAR PHYSICS PARAMETERS²⁴

| Parameter | Symbol | Value | Units |
|--|-----------------------|-----------------------|--------------------------------------|
| Total mass | M⊙ | 1.99×10^{33} | g |
| Radius | R_{\odot} | 6.96×10^{10} | cm |
| Surface gravity | g⊙ | 2.74×10^4 | $cm s^{-2}$ |
| Escape speed | v_{∞} | 6.18×10^{7} | $cm s^{-1}$ |
| Upward mass flux in spicules | | 1.6×10^{-9} | $g cm^{-2} s^{-1}$ |
| Vertically integrated atmospheric density | _ | 4.28 | g cm ⁻² |
| Sunspot magnetic field strength | B_{\max} | 2500-3500 | G |
| Surface temperature | T_0 | 6420 | K |
| Radiant power | \mathcal{L}_{\odot} | 3.90×10^{33} | erg s ⁻¹ |
| Radiant flux density | F | 6.41×10^{10} | erg cm ⁻² s ⁻¹ |
| Optical depth at 500 nm, measured from photosphere | 7 ₅₀₀ | 0.99 | _ |
| Astronomical unit (radius of earth's orbit) | AU | 1.50×10^{13} | cm |
| Solar constant (intensity at 1 AU) | f | 1.39×10^6 | $erg cm^{-2} s^{-1}$ |

Chromosphere and Corona²⁵

| Parameter (Units) | Quiet Šun | Coronal Hole | Active Region |
|--|------------------------------|---------------------|----------------------------------|
| Chromospheric radiation losses (erg cm ⁻² s ⁻¹) | | | |
| Low chromosphere | 2×10^6 | 2×10^6 | $\gtrsim 10^7$ |
| Middle chromosphere | 2×10^6 | 2×10^6 | 10 ⁷ |
| Upper chromosphere | 3×10^5 | 3×10^5 | 2×10^6 |
| Total | 4×10^6 | 4×10^6 | $\gtrsim 2 \times 10^7$ |
| Transition layer pressure (dyne cm ⁻²) | 0.2 | 0.07 | 2 |
| Coronal temperature (K) at 1.1 R _O | $1.1-1.6 \times 10^6$ | 10 ⁶ | 2.5×10^{6} |
| Coronal energy losses (erg cm ⁻² s ⁻¹) | | | |
| Conduction | 2×10^5 | 6 × 10 ⁴ | 10 ⁵ -10 ⁷ |
| Radiation | 10 ⁵ | 104 | 5×10^6 |
| Solar Wind | $\lesssim 5 \times 10^4$ | 7×10^5 | < 10 ⁵ |
| Total | 3×10^5 | 8×10^5 | 10 ⁷ |
| Solar wind mass loss (g cm ⁻² s ⁻¹) | $\lesssim 2 \times 10^{-11}$ | 2×10^{-10} | $< 4 \times 10^{-11}$ |

THERMONUCLEAR FUSION²⁶

Natural abundance of isotopes:

hydrogen
$$n_D/n_H = 1.5 \times 10^{-4}$$

helium $n_{\text{He}^3}/n_{\text{He}^4} = 1.3 \times 10^{-6}$
lithium $n_{\text{Li}^6}/n_{\text{Li}^7} = 0.08$

Mass ratios:

$$m_e/m_D$$
 = 2.72 × 10⁻⁴ = 1/3670
 $(m_e/m_D)^{1/2}$ = 1.65 × 10⁻² = 1/60.6
 m_e/m_T = 1.82 × 10⁻⁴ = 1/5496
 $(m_e/m_T)^{1/2}$ = 1.35 × 10⁻² = 1/74.1

Absorbed radiation dose is measured in rads: 1 rad = $10^2 \,\mathrm{erg}\,\mathrm{g}^{-1}$. The curie (abbreviated Ci) is a measure of radioactivity: 1 curie = 3.7 x 10^{10} counts sec⁻¹.

Fusion reactions (branching ratios are correct for energies near the cross section peaks; a negative yield means the reaction is endothermic):27

(1a) D + D
$$\rightarrow T(1.01 \text{ MeV}) + p(3.02 \text{ MeV})$$

(1b) $\rightarrow He^3(0.82 \text{ MeV}) + n(2.45 \text{ MeV})$
(2) D + T $\rightarrow He^4(3.5 \text{ MeV}) + n(14.1 \text{ MeV})$

(2)
$$D + T \longrightarrow He^4(3.5 \text{ MeV}) + n(14.1 \text{ MeV})$$

(3)
$$D + He^3 \longrightarrow He^4 (3.6 \text{ MeV}) + p(14.7 \text{ MeV})$$

(4)
$$T + T \longrightarrow He^4 + 2n + 11.3 MeV$$

(5a)
$$He^3 + T \xrightarrow{51\%} He^4 + p + n + 12.1 MeV$$

(5b)
$$+ D(9.5 \text{ MeV})$$

(5c)
$$\xrightarrow{6\%} \text{He}^{5}(2.4 \,\text{MeV}) + p(11.9 \,\text{MeV})$$

(6) $p + \text{Li}^{6} \xrightarrow{6\%} \text{He}^{4}(1.7 \,\text{MeV}) + \text{He}^{3}(2.3 \,\text{MeV})$

(6)
$$p + Li^6 \longrightarrow He^4(1.7 \text{ MeV}) + He^3(2.3 \text{ MeV})$$

(7a)
$$p + Li^7 \xrightarrow{20\%} 2 He^4 + 17.3 MeV$$

(7b)
$$\frac{20\%}{80\%} \text{Be}^7 + \text{n} - 1.6 \text{ MeV}$$

(8) D + Li⁶ $\xrightarrow{2} \text{2He}^4 + 22.4 \text{ MeV}$

$$(8) D + Li6 \longrightarrow 2He4 + 22.4 MeV$$

(9)
$$p + B^{11} \longrightarrow 3 He^4 + 8.7 MeV$$

(10)
$$n + Li^6 \longrightarrow He^4(2.1 MeV) + T(2.7 MeV)$$

The total cross section in barns as a function of E, the energy in keV of the incident particle [the first ion on the left side of Eqs. (1)-(5)], assuming the target ion at rest, can be fitted by 28

$$\sigma_T(E) = \frac{A_5 + \left[(A_4 - A_3 E)^2 + 1 \right]^{-1} A_2}{E \left[\exp(A_1 E^{-1/2}) - 1 \right]}$$

where the Duane coefficients A_j for the principle fusion reactions are as follows:

| | D-D (1a) | D-D (1b) | D-T (2) | D-He ³ (3) | T-T (4) | T-He ³ (5a-c) |
|---------|-----------------------|-----------------------|------------------------|-----------------------|---------------------|-----------------------------|
| A_1 | 46.097 | 47.88 | 45.95 | 89.27 | 38.39 | 123.1 |
| A_2 | 372 | 482 | 50200 | 25900 | 448 | 11250 |
| A_3 | 4.36×10^{-4} | 3.08×10^{-4} | 1.368×10^{-2} | 3.98×10^{-3} | 1.02×10^{-3} | 0 |
| A_{4} | 1.220 | 1.177 | 1.076 | 1.297 | 2.09 | 0 |
| A_5 | 0 | 0 | 409 | 647 | 0 | 0 |

Reaction rates $\overline{\sigma v}$ (in cm³ sec⁻¹), averaged over Maxwellian distributions:

| Temperature (keV) | D-D (la + lb) | D-T (2) | D-He ³ (3) | T-T (4) | T-He ³ (5a-c) |
|-------------------|-----------------------|-----------------------|-----------------------|-----------------------|-----------------------------|
| 1.0 | 1.5×10^{-22} | 5.5×10^{-21} | 10-26 | 3.3×10^{-22} | 10-28 |
| 2.0 | 5.4×10^{-21} | 2.6×10^{-19} | 1.4×10^{-23} | 7.1×10^{-21} | 10-25 |
| 5.0 | 1.8×10^{-19} | 1.3×10^{-17} | 6.7×10^{-21} | 1.4×10^{-19} | 2.1×10^{-22} |
| 10.0 | 1.2×10^{-18} | 1.1×10^{-16} | 2.3×10^{-19} | 7.2×10^{-19} | 1.2×10^{-20} |
| 20.0 | 5.2×10^{-18} | 4.2×10^{-16} | 3.8×10^{-18} | 2.5×10^{-18} | 2.6×10^{-19} |
| 50.0 | 2.1×10^{-17} | 8.7×10^{-16} | 5.4×10^{-17} | 8.7×10^{-18} | 5.3×10^{-18} |
| 100.0 | 4.5×10^{-17} | 8.5×10^{-16} | 1.6×10^{-16} | 1.9×10^{-17} | 2.7×10^{-17} |
| 200.0 | 8.8×10^{-17} | 6.3×10^{-16} | 2.4×10^{-16} | 4.2×10^{-17} | 9.2×10^{-17} |
| 500.0 | 1.8×10^{-16} | 3.7×10^{-16} | 2.3×10^{-16} | 8.4×10^{-17} | 2.9×10^{-16} |
| 1000.0 | 2.2×10^{-16} | 2.7×10^{-16} | 1.8×10^{-16} | 8.0×10^{-17} | 5.2×10^{-16} |

For low energies ($T\lesssim 25\,\mathrm{keV}$) the data may be represented by

$$(\overline{\sigma v})_{DD} = 2.33 \times 10^{-14} T^{-2/3} \exp(-18.76 T^{-1/3}) \text{ cm}^3 \text{ sec}^{-1};$$

$$(\overline{\sigma v})_{DT} = 3.68 \times 10^{-12} T^{-2/3} \exp(-19.94 T^{-1/3}) \text{ cm}^3 \text{ sec}^{-1},$$

where T is measured in keV.

The power density released in the form of charged particles is

$$P_{DD} = 3.3 \times 10^{-13} n_D^2 (\overline{\sigma v})_{DD}$$
 watt cm⁻³ (including the subsequent D-T reaction);

$$P_{DT} = 5.6 \times 10^{-13} n_D n_T (\overline{\sigma v})_{DT} \text{ watt cm}^{-3};$$

$$P_{D{
m He}^3} = 2.9 \times 10^{-12} n_D n_{{
m He}^3} (\overline{\sigma v})_{D{
m He}^3} \, {
m watt} \, {
m cm}^{-3}.$$

RELATIVISTIC ELECTRON BEAMS

Here $\gamma = (1 - \beta^2)^{-1/2}$ is the relativistic scaling factor; quantities in analytic formulas are expressed in SI or cgs units, as indicated: in numerical formulas, I is in amperes (A), B is in gauss (G), electron linear density N is in cm⁻¹, and temperature, voltage and energy are in MeV; $\beta_z = v_z/c$: k is Boltzmann's constant.

Relativistic electron gyroradius:

$$r_e = \frac{mc^2}{eB} (\gamma^2 - 1)^{1/2} (\text{cgs}) = 1.70 \times 10^3 (\gamma^2 - 1)^{1/2} B^{-1} \text{ cm}.$$

Relativistic electron energy:

$$W = mc^2 \gamma = 0.511 \gamma \text{ MeV}.$$

Bennett pinch condition:

$$I^2 = 2Nk(T_e + T_i)c^2 \text{ (cgs)} = 3.20 \times 10^{-4}N(T_e + T_i) \text{ A}^2.$$

Alfvén-Lawson limit:

$$I_A = (mc^3/e)\beta_z\gamma (cgs) = (4\pi mc/\mu_0 e)\beta_z\gamma (SI) = 1.70 \times 10^4\beta_z\gamma A.$$

The ratio of net current to I_A is

$$\frac{I}{I_A} = \frac{\nu}{\gamma}.$$

Here $\nu = Nr_e$ is the Budker number, where $r_e = e^2/mc^2 = 2.82 \times 10^{-13}$ cm is the classical electron radius. Beam electron number density is

$$n_b = 2.08 \times 10^8 J \beta^{-1} \text{ cm}^{-3}$$

where J is the current density in $A \text{ cm}^{-2}$. For a uniform beam of radius a (in cm),

$$n_b = 6.63 \times 10^7 Ia^{-2} \beta^{-1} \text{ cm}^{-3}$$

and

$$\frac{2r_e}{a}=\frac{\nu}{\gamma}.$$

Child's law: (non-relativistic) space-charge-limited current density between parallel plates with voltage drop V and separation d (in cm) is

$$J = 2.34 \times 10^3 V^{3/2} d^{-2} \,\mathrm{A \, cm^{-2}}.$$

The saturated parapotential current (magnetically self-limited flow along equipotentials in pinched diodes and transmission lines) is 29

$$I_p = 8.5 \times 10^3 G \gamma \ln \left[\gamma + (\gamma^2 - 1)^{1/2} \right] A.$$

where G is a geometrical factor depending on the diode structure:

$$G = \frac{w}{2\pi d}$$
 for parallel plane cathode and anode of width w , separation d ; $G = \left(\ln \frac{R_2}{R_1}\right)^{-1}$ for cylinders of radii R_1 (inner) and R_2 (outer); $G = \frac{R_c}{d_0}$ for conical cathode of radius R_c , maximum separation d_0 (at $r = R_c$) from plane anode.

For $\beta \to 0 \ (\gamma \to 1)$, both I_A and I_p vanish.

The condition for a longitudinal magnetic field B_z to suppress filamentation in a beam of current density J (in $A cm^{-2}$) is

$$B_z > 47\beta_z (\gamma J)^{1/2} G.$$

Voltage registered by Rogowski coil of minor cross-sectional area A, n turns, major radius a, inductance L, external resistance R and capacitance C (all in SI):

externally integrated
$$V = (1/RC)(nA\mu_0I/2\pi a);$$

self-integrating $V = (R/L)(nA\mu_0I/2\pi a) = RI/n.$

X-ray production, target with average atomic number Z ($V \lesssim 5 \,\text{MeV}$):

$$\eta \equiv x$$
-ray power/beam power = $7 \times 10^{-4} ZV$.

X-ray dose at 1 meter generated by an e-beam depositing total charge Q coulombs while $V \ge 0.84V_{\text{max}}$ in material with charge state Z:

$$D = 150 V_{\text{max}}^{2.8} Q Z^{1/2} \text{ rads.}$$

BEAM INSTABILITIES30

| Name | Conditions | Saturation Mechanism |
|---------------------------------------|--|---|
| Electron- electron | $V_d > \tilde{V}_{ej}, j = 1.2$ | Electron trapping until $\bar{V}_{ej} \sim V_d$ |
| Buneman | $V_d > (M/m)^{1/3} \bar{V}_i,$ $V_d > \bar{V}_e$ | Electron trapping until $ar{V}_e \sim V_d$ |
| Beam-plasma | $V_b > (n_p/n_b)^{1/3} \bar{V}_b$ | Trapping of beam electrons |
| Weak beam- plasma | $V_b < (n_p/n_b)^{1/3} \bar{V}_b$ | Quasilinear or nonlinear (mode coupling) |
| Beam-plasma (hot-electron) | $ar{V}_{e} > V_{b} > ar{V}_{b}$ | Quasilinear or nonlinear |
| Ion acoustic | $T_e \gg T_i, \ V_d \gg C_{\bullet}$ | Quasilinear, ion tail formation, nonlinear scattering, or resonance broadening. |
| Anisotropic temperature (hydro) | $T_{e\perp} > 2T_{e\parallel}$ | Isotropization |
| Ion cyclotron | $V_d > 20 \bar{V}_i \; 	ext{(for} \ T_e pprox T_i 	ext{)}$ | Ion heating |
| Beam-cyclotron (hydro) | $V_d > C_s$ | Resonance broadening |
| Modified two- stream (hydro) | $V_d < (1+\beta)^{1/2} V_A,$ $V_d > C_a$ | Trapping |
| Ion-ion (equal beams) | $U < 2(1+\beta)^{1/2} V_A$ | Ion trapping |
| Ion-ion (equal beams) | $U < 2C_s$ | Ion trapping |

For nomenclature, see p. 50.

| | Paramet | ers of Most Unsta | able Mode | |
|---------------------------------------|---|--|----------------------------|-------------------------------------|
| Name | Growth Rate | Frequency | Wave Number | Group Velocity |
| Electron- electron | $\frac{1}{2}\omega_e$ | 0 | $0.9 \frac{\omega_e}{V_d}$ | 0 |
| Buneman | $0.7 \left(\frac{m}{M}\right)^{1/3} \omega_e$ | $0.4 \left(\frac{m}{M}\right)^{1/3} \omega_e$ | $\frac{\omega_e}{V_d}$ | $\frac{2}{3}V_d$ |
| Beam-plasma | $0.7 \left(\frac{n_b}{n_p}\right)^{1/3} \omega_e$ | $\left \frac{\omega_e - \left(\frac{n_b}{n_b} \right)^{1/3} \omega_e}{\omega_e} \right $ | $\frac{\omega_e}{V_b}$ | $\frac{2}{3}V_b$ |
| Weak beam- plasma | $\left \frac{n_b}{2n_p} \left(\frac{V_b}{\bar{V}_b} \right)^2 \omega_e \right $ | ω_e | $\frac{\omega_e}{V_b}$ | $\frac{3\bar{V}_e^2}{V_b}$ |
| Beam-plasma (hot-electron) | $\left(\frac{n_b}{n_p}\right)^{1/2} \frac{\bar{V}_e}{V_b} \omega_e$ | $rac{V_b}{ar{V}_e}\omega_e$ | λ_D^{-1} | V_{b} |
| Ion acoustic | $\left(\frac{m}{M}\right)^{1/2}\omega_i$ | ω_i | λ_D^{-1} | C. |
| Anisotropic temperature (hydro) | Ωε | $\omega_e \cos \theta \sim \Omega_e$ | r_e^{-1} | $ar{V}_{e\perp}$ |
| Ion cyclotron | $0.1\Omega_i$ | $1.2\Omega_i$ | r_i^{-1} | $\frac{1}{3}ar{V}_i$ |
| Beam-cyclotron (hydro) | 0.7Ωε | $n\Omega_e$ | $0.7\lambda_D^{-1}$ | $\gtrsim V_d; \lesssim C_{\bullet}$ |
| Modified two- stream (hydro) | $rac{1}{2}\Omega_{H}$ | $0.9\Omega_{H}$ | $1.7 \frac{\Omega_H}{V_d}$ | |
| Ion-ion (equal beams) | $0.4\Omega_{H}$ | 0 | $1.2\frac{\Omega_H}{U}$ | 0 |
| Ion-ion (equal beams) | $0.4\omega_i$ | 0 | $1.2 \frac{\omega_i}{U}$ | 0 |

For nomenclature, see p. 50.

In the preceding tables, subscripts e, i, d, b, p stand for "electron," "ion," "drift," "beam," and "plasma," respectively. Thermal velocities are denoted by a bar. In addition, the following are used:

| m | electron mass | r_e , r_i | gyroradius |
|---------------------------------|------------------|--|----------------------------------|
| M | ion mass | 3 | plasma/magnetic energy |
| V | velocity | | density ratio |
| T | temperature | $V_{\mathcal{A}}$ | Alfvén speed |
| n_e, n_i | number density | $\Omega_{\epsilon}, \Omega_{\epsilon}$ | gyrofrequency |
| n | harmonic number | Ω_{H} | hybrid gyrofrequency, |
| $C_{\bullet} = (T_{e}/M)^{1/2}$ | ion sound speed | | $\Omega_H^2 = \Omega_e \Omega_i$ |
| ω_e, ω_i | plasma frequency | $oldsymbol{U}$ | relative drift velocity of |
| $\lambda_{\mathcal{D}}$ | Debye length | | two ion species |

LASERS

System Parameters

Efficiencies and power levels are approximately state-of-the-art (1987).31

| Type | Wavelength | Efficiency | Power levels ava | ilable (W) |
|-----------------|------------|-----------------------|----------------------------------|---------------------|
| Турс | (μm) | Linciency | Pulsed | CW |
| CO ₂ | 10.6 | 0.01-0.02 (pulsed) | > 2 × 10 ¹³ | > 10 ⁵ |
| CO | 5 | 0.4 | > 10 ⁹ | > 100 |
| Holmium | 2.06 | 0.03 | > 10 ⁷ | 30 |
| Iodine | 1.315 | 0.003 | > 10 ¹² | - |
| Nd-glass, | 1.06 | 0.001-0.06 | ~ 10 ¹⁴ | 1-10 ³ |
| YAG | | | (10-beam system) | |
| *Color | 1-4 | 10-3 | > 10 ⁶ | 1 |
| center | | | | |
| *OPO | 0.7-0.9 | 10-3 | 10 ⁶ | 1 |
| Ruby | 0.6943 | $< 10^{-3}$ | 1010 | 1 |
| He-Ne | 0.6328 | 10-4 | - | $1-50\times10^{-3}$ |
| *Argon ion | 0.45-0.60 | 10-3 | 5×10^4 | 1-10 |
| N ₂ | 0.3371 | 0.001-0.05 | 10 ⁵ -10 ⁶ | - |
| *Dye | 0.3-1.1 | 10-3 | > 10 ⁶ | 140 |
| Kr-F | 0.26 | 0.08 | > 10° | _ |
| Xenon | 0.175 | 0.02 | > 10 ⁸ | - |

^{*}Tunable sources

YAG stands for Yttrium-Aluminum Garnet and OPO for Optical Parametric Oscillator.

Formulas

An e-m wave with k | B has an index of refraction given by

$$n_{\pm} = \left[1 - \omega_{ne}^2/\omega(\omega \mp \omega_{ce})\right]^{1/2},$$

where \pm refers to the helicity. The rate of change of polarization angle θ as a function of displacement s (Faraday rotation) is given by

$$d\theta/ds = (k/2)(n_- - n_+) = 2.36 \times 10^4 NBf^{-2} cm^{-1}$$

where N is the electron number density, B is the field strength, and f is the wave frequency, all in cgs.

The quiver velocity of an electron in an e-m field of angular frequency ω is

$$v_0 = eE_{\rm max}/m\omega = 25.6I^{1/2}\lambda_0 \,{\rm cm\,sec}^{-1}$$

in terms of the laser flux $I=cE_{\rm max}^2/8\pi$, with I in watt/cm², laser wavelength λ_0 in μ m. The ratio of quiver energy to thermal energy is

$$W_{\rm qu}/W_{\rm th} = m_e v_0^2/2kT = 1.81 \times 10^{-13} \lambda_0^2 I/T$$

where T is given in eV. For example, if $I = 10^{15} \, \mathrm{W \, cm^{-2}}$, $\lambda_0 = 1 \, \mu \mathrm{m}$, $T = 2 \, \mathrm{keV}$, then $W_{\mathrm{qu}}/W_{\mathrm{th}} \approx 0.1$.

Pondermotive force:

$$\mathcal{F} = N\nabla \langle E^2 \rangle / 8\pi N_c,$$

where

$$N_c = 1.1 \times 10^{21} \lambda_0^{-2} \text{cm}^{-3}$$
.

For uniform illumination of a lens with f-number F, the diameter d at focus (85% of the energy) and the depth of focus l (distance to first zero in intensity) are given by

$$d \approx 2.44 F \lambda \theta / \theta_{DL}$$
 and $l \approx \pm 2 F^2 \lambda \theta / \theta_{DL}$.

Here θ is the beam divergence containing 85% of energy and θ_{DL} is the diffraction-limited divergence:

$$\theta_{DL} = 2.44 \lambda/b$$

where b is the aperture. These formulas are modified for nonuniform (such as Gaussian) illumination of the lens or for pathological laser profiles.

ATOMIC PHYSICS AND RADIATION

Energies and temperatures are in eV; all other units are cgs except where noted. Z is the charge state (Z=0 refers to a neutral atom); the subscript e labels electrons. N refers to number density, n to principal quantum number. Asterisk superscripts on level population densities denote local thermodynamic equilibrium (LTE) values. Thus N_n^* is the LTE number density of atoms (or ions) in level n.

Characteristic atomic collision cross section:

(1)
$$\pi a_0^2 = 8.80 \times 10^{-17} \,\mathrm{cm}^2$$
.

Binding energy of outer electron in level labelled by quantum numbers n, l:

(2)
$$E_{\infty}^{Z}(n,l) = -\frac{Z^{2}E_{\infty}^{H}}{(n-\Delta_{l})^{2}},$$

where $E_{\infty}^{H}=13.6\,\mathrm{eV}$ is the hydrogen ionization energy and $\Delta_{l}=0.75l^{-5}$, $l\gtrsim 5$, is the quantum defect.

Excitation and Decay

Cross section (Bethe approximation) for electron excitation by dipole allowed transition $m \to n$ (Refs. 32, 33):

(3)
$$\sigma_{mn} = 2.36 \times 10^{-13} \frac{f_{nm}g(n,m)}{\epsilon \Delta E_{nm}} \text{ cm}^2,$$

where f_{nm} is the oscillator strength, g(n, m) is the Gaunt factor, ϵ is the incident electron energy, and $\Delta E_{nm} = E_n - E_m$.

Electron excitation rate averaged over Maxwellian velocity distribution, $X_{mn} = N_e \langle \sigma_{mn} v \rangle$ (Refs. 34, 35):

(4)
$$X_{mn} = 1.6 \times 10^{-5} \frac{f_{nm} \langle g(n,m) \rangle N_e}{\Delta E_{nm} T_e^{1/2}} \exp\left(-\frac{\Delta E_{nm}}{T_e}\right) \sec^{-1},$$

where (g(n, m)) denotes the thermal averaged Gaunt factor (generally ~ 1 for atoms, ~ 0.2 for ions).

Rate for electron collisional deexcitation:

(5)
$$Y_{nm} = (N_m^*/N_n^*)X_{mn}.$$

Here $N_m^*/N_n^* = (g_m/g_n) \exp(\Delta E_{nm}/T_e)$ is the Boltzmann relation for level population densities, where g_n is the statistical weight of level n.

Rate for spontaneous decay $n \rightarrow m$ (Einstein A coefficient)³⁴

(6)
$$A_{nm} = 4.3 \times 10^{7} (g_n/g_m) f_{nm} (\Delta E_{nm})^2 \sec^{-1}.$$

Intensity emitted per unit volume from the transition $n \rightarrow m$ in an optically thin plasma:

(7)
$$I_{nm} = 1.6 \times 10^{-19} A_{nm} N_n \Delta E_{nm} \text{ watt/cm}^3.$$

Condition for steady state in a corona model:

$$(8) N_0 N_e \langle \sigma_{0n} v \rangle = N_n A_{n0},$$

where the ground state is labelled by a zero subscript.

Hence for a transition $n \to m$ in ions, where $\langle g(n,0) \rangle \approx 0.2$.

(9)
$$I_{nm} = 5.1 \times 10^{-25} \frac{f_{nm} g_0 N_e N_0}{g_m T_e^{1/2}} \left(\frac{\Delta E_{nm}}{\Delta E_{n0}}\right)^3 \exp\left(-\frac{\Delta E_{n0}}{T_e}\right) \frac{\text{watt}}{\text{cm}^3}.$$

Ionization and Recombination

In a general time-dependent situation the number density of the charge state Z satisfies

(10)
$$\frac{dN(Z)}{dt} = N_{\epsilon} \left[-S(Z)N(Z) - \alpha(Z)N(Z) + S(Z-1)N(Z-1) + \alpha(Z+1)N(Z+1) \right].$$

Here S(Z) is the ionization rate. The recombination rate $\alpha(Z)$ has the form $\alpha(Z) = \alpha_r(Z) + N_e \alpha_3(Z)$, where α_r and α_3 are the radiative and three-body recombination rates, respectively:

Classical ionization cross-section 36 for any atomic shell j

(11)
$$\sigma_i = 6 \times 10^{-14} b_j g_j(x) / U_j^2 \text{ cm}^2.$$

Here b_j is the number of shell electrons: U_j is the binding energy of the ejected electron: $x = \epsilon/U_j$, where ϵ is the incident electron energy: and g is a universal function with a minimum value $g_{\min} \approx 0.2$ at $x \approx 4$.

Ionization from ion ground state, averaged over Maxwellian electron distribution, for $0.02 \lesssim T_e/E_\infty^Z \lesssim 100$ (Ref. 35):

(12)
$$S(Z) = 10^{-5} \frac{(T_e/E_{\infty}^Z)^{1/2}}{(E_{\infty}^Z)^{3/2}(6.0 + T_e/E_{\infty}^Z)} \exp\left(-\frac{E_{\infty}^Z}{T_e}\right) \text{ cm}^3/\text{sec.}$$

where E_{∞}^{Z} is the ionization energy.

Electron-ion radiative recombination rate $(e + N(Z) \rightarrow N(Z - 1) + h\nu)$ for $T_e/Z^2 \lesssim 400 \,\text{eV}$ (Ref. 37):

(13)
$$\alpha_r(Z) = 5.2 \times 10^{-14} Z \left(\frac{E_{\infty}^Z}{T_e}\right)^{1/2} \left[0.43 + \frac{1}{2} \ln(E_{\infty}^Z/T_e) + 0.469 (E_{\infty}^Z/T_e)^{-1/3}\right] \text{cm}^3/\text{sec.}$$

For $1 \, \text{eV} < T_e/Z^2 < 15 \, \text{eV}$, this becomes approximately 35

(14)
$$\alpha_r(Z) = 2.7 \times 10^{-13} Z^2 T_e^{-1/2} \text{ cm}^3/\text{sec.}$$

Collisional (three-body) recombination rate for singly ionized plasma: 38

(15)
$$\alpha_3 = 8.75 \times 10^{-27} T_e^{-4.5} \text{ cm}^6/\text{sec.}$$

Photoionization cross section for ions in level n, l (short-wavelength limit):

(16)
$$\sigma_{\rm ph}(n,l) = 1.64 \times 10^{-16} Z^5 / n^3 K^{7+2l} \, {\rm cm}^2,$$

where K is the wavenumber in Rydbergs (1 Rydberg = 1.0974×10^5 cm⁻¹).

Ionization Equilibrium Models

Saha equilibrium:39

(17)
$$\frac{N_e N_1^*(Z)}{N_n^*(Z-1)} = 6.0 \times 10^{21} \frac{g_1^Z T_e^{3/2}}{g_n^{Z-1}} \exp\left(-\frac{E_{\infty}^Z(n,l)}{T_e}\right) \text{ cm}^{-3},$$

where g_n^Z is the statistical weight for level n of charge state Z and $E_{\infty}^Z(n,l)$ is the ionization energy of the neutral atom initially in level (n,l), given by Eq. (2).

In a steady state at high electron density,

(18)
$$\frac{N_e N^*(Z)}{N^*(Z-1)} = \frac{S(Z-1)}{\alpha_3},$$

a function only of T.

Conditions for LTE:39

(a) Collisional and radiative excitation rates for a level n must satisfy

$$(19) Y_{nm} \gtrsim 10A_{nm}.$$

(b) Electron density must satisfy

(20)
$$N_e \gtrsim 7 \times 10^{18} Z^7 n^{-17/2} (T/E_{\infty}^Z)^{1/2} \text{cm}^{-3}.$$

Steady state condition in corona model:

(21)
$$\frac{N(Z-1)}{N(Z)} = \frac{\alpha_r}{S(Z-1)}.$$

Corona model is applicable if 40

(22)
$$10^{12}t_I^{-1} < N_e < 10^{16}T_e^{7/2} \,\mathrm{cm}^{-3}.$$

where t_I is the ionization time.

Radiation

N. B. Energies and temperatures are in eV: all other quantities are in cgs units except where noted. Z is the charge state (Z=0 refers to a neutral atom); the subscript e labels electrons. N is number density.

Average radiative decay rate of a state with principal quantum number n is

(23)
$$A_n = \sum_{m \le n} A_{nm} = 1.6 \times 10^{10} Z^4 n^{-9/2} \text{ sec.}$$

Natural linewidth (ΔE in eV):

(24)
$$\Delta E \Delta t = h = 4.14 \times 10^{-15} \text{ eV sec.}$$

where Δt is the lifetime of the line.

Doppler width:

(25)
$$\Delta \lambda / \lambda = 7.7 \times 10^{-5} (T/\mu)^{1/2},$$

where μ is the mass of the emitting atom or ion scaled by the proton mass. Optical depth for a Doppler-broadened line:³⁹

(26)
$$\tau = 1.76 \times 10^{-13} \lambda (Mc^2/kT)^{1/2} NL = 5.4 \times 10^{-9} \lambda (\mu/T)^{1/2} NL.$$

where λ is the wavelength and L is the physical depth of the plasma; M, N, and T are the mass, number density, and temperature of the absorber; μ is M divided by the proton mass. Optically thin means $\tau < 1$.

Resonance absorption cross section at center of line:

(27)
$$\sigma_{\lambda=\lambda_c} = 5.6 \times 10^{-13} \lambda^2 / \Delta \lambda \text{ cm}^2.$$

Wien displacement law (wavelength of maximum black-body emission):

(28)
$$\lambda_{\text{max}} = 2.50 \times 10^{-5} T^{-1} \text{ cm}.$$

Radiation from the surface of a black body at temperature T:

(29)
$$W = 1.03 \times 10^5 T^4 \text{ watt/cm}^2$$
.

Bremsstrahlung from hydrogen-like plasma:26

(30)
$$P_{\rm Br} = 1.69 \times 10^{-32} N_e T_e^{1/2} \sum \left[Z^2 N(Z) \right] \text{ watt/cm}^3,$$

where the sum is over all ionization states Z.

Bremsstrahlung optical depth:41

(31)
$$\tau = 5.0 \times 10^{-38} N_e N_i Z^2 \overline{g} L T^{-7/2},$$

where $\overline{g} \approx 1.2$ is an average Gaunt factor and L is the physical path length. Inverse bremsstrahlung absorption coefficient⁴² for radiation of angular frequency ω :

(32)
$$\kappa = 3.1 \times 10^{-7} Z n_e^2 \ln \Lambda T^{-3/2} \omega^{-2} (1 - \omega_p^2 / \omega^2)^{1/2} \text{ cm}^{-1};$$

here Λ is the electron thermal velocity divided by V, where V is the larger of ω and ω_p multiplied by the larger of Ze^2/kT and $\hbar/(mkT)^{1/2}$.

Recombination (free-bound) radiation:

(33)
$$P_r = 1.69 \times 10^{-32} N_e T_e^{1/2} \sum_{e} \left[Z^2 N(Z) \left(\frac{E_{\infty}^{Z-1}}{T_e} \right) \right] \text{ watt/cm}^3.$$

Cyclotron radiation²⁶ in magnetic field B:

(34)
$$P_c = 6.21 \times 10^{-28} B^2 N_e T_e \text{ watt/cm}^3.$$

For $N_e kT_e = N_i kT_i = B^2/16\pi$ ($\beta = 1$, isothermal plasma),²⁶

(35)
$$P_c = 5.00 \times 10^{-38} N_e^2 T_e^2 \text{ watt/cm}^3.$$

Cyclotron radiation energy loss e-folding time for a single electron:⁴¹

(36)
$$t_c \approx \frac{9.0 \times 10^8 B^{-2}}{2.5 + \gamma} \text{ sec},$$

where γ is the kinetic plus rest energy divided by the rest energy mc^2 . Number of cyclotron harmonics⁴¹ trapped in a medium of finite depth L:

(37)
$$m_{\rm tr} = (57\beta BL)^{1/6},$$

where $\beta = 8\pi NkT/B^2$.

Line radiation is given by summing Eq. (9) over all species in the plasma.

REFERENCES

Most of the formulas and data in this collection are well known and for all practical purposes are in the "public domain." The books and articles cited below are intended primarily not for the purpose of giving credit to the original workers, but (1) to guide the reader to sources containing related material and (2) to indicate where to find derivations, explanations, examples, etc., which have been omitted from this compilation. Additional material can also be found in D. L. Book, NRL Memorandum Report No. 3332 (1977).

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